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Influence of redox processes on the germanium isotopic composition
of ordinary chondrites

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Abstract

Ordinary chondrites (OCs) are classified into three groups, according to their oxidation state, which increases from the H to L to LL groups. This is demonstrated by the decrease in metal content (H=\(~8\) vol\%, L=\(~4\) vol\%, and LL=\(~2\) vol\%), and by a positive correlation between \(\Delta^{17}\text{O}\) and \%Fa through the OC sequence. Compared to other chondrites, OCs exhibit the largest variation in oxidation state but there is an ongoing debate on the processes that control this variation. To constrain the causes of the variations in the oxidation state with respect to the associated nebular versus parent bodies processes, we investigated the elemental and isotopic
variations of germanium (moderately siderophile and volatile) in the bulk sample, as well as in
the metal, silicate and sulfide phases, over a range of petrographic types for the H, L, and LL
ordinary chondrites.

We found that $\delta^{74/70}{\text{Ge}}_{\text{metal}}$ is a proxy for the $\delta^{74/70}{\text{Ge}}_{\text{bulk}}$ composition and that each OC group is
distinguishable by their $\delta^{74/70}{\text{Ge}}_{\text{metal}}$, which increases from $-0.51\pm0.09\%\text{o}$ for H chondrites, $-0.31\pm0.16\%\text{o}$ for L chondrites, and, finally, to $-0.26\pm0.09\%\text{o}$ for LL chondrites (2σSD). Additionally, the OC sequence exhibited a positive correlation, from H to L to LL, between $\delta^{74/70}{\text{Ge}}_{\text{metal}}$ and %Fa, as well as oxygen isotopes ($\delta^{17}{\text{O}}, \delta^{18}{\text{O}}$ and $\Delta^{17}{\text{O}}$), that was not a consequence of a "size sorting effect" on chondrules (i.e. chondrule mixing) or metamorphic processes in the parent bodies but, rather, was the result of nebular processes. We propose that the correlation between the $\delta^{74/70}{\text{Ge}}$ values and %Fa, $\Delta^{17}{\text{O}}, \delta^{18}{\text{O}}$ can be explained by an increasing proportion of accreted hydrated phyllosilicates, from the H, L to LL groups, with high $\delta^{74/70}{\text{Ge}}$ and $\Delta^{17}{\text{O}}$. We found that 10 to 15% of phyllosilicates, with a composition of $[\text{Ge}]=4-7$ ppm and $\delta^{74/70}{\text{Ge}}=3-2.5\%\text{o}$, is needed to change the $\delta^{74/70}{\text{Ge}}$ from H to LL, which corresponds to a $\Delta^{17}{\text{O}}=6.7\%\text{o}$. This value agrees with the $\Delta^{17}{\text{O}}=7\%\text{o}$ composition of the accreted nebular component reported by Choi et al. (1998). During thermal metamorphism, phyllosilicates destabilize, liberating germanium that will be incorporated in the metal, then leading to its high $\delta^{74/70}{\text{Ge}}$ signature.

High-temperature metamorphism can explain the lack of $\delta^{74/70}{\text{Ge}}_{\text{metal}}$ variation with the petrologic
type in the OC, even for the type 3 chondrites ($T\approx675{\degree}\text{C}$), implying a complete reaction even at
low petrologic types. In addition, metal-silicate re-equilibration in response to thermal
metamorphism results in a decrease in $\Delta^{74/70}{\text{Ge}}_{\text{metal-silicate}}$ from 0.33‰ to 0.06‰, within the H
chondrite group, which is interpreted as the result of $\delta^{74/70}{\text{Ge}}_{\text{silicate}}$ variation. The mean positive
$\Delta^{74/70}{\text{Ge}}_{\text{metal-silicate}}$ fractionation factor of $+0.22\pm0.36\%\text{o}$ (error propagation on individual error) also
displays a remarkable similarity to the direction of isotopic fractionation with other germanium
isotopic metal-silicate datasets, such as the magmatic iron meteorites, the Earth silicate reservoirs. We propose that the $\Delta^{74/70}\text{Ge}_{\text{metal-silicate}}$ and the negative $\Delta^{74/70}\text{Ge}$ values of OCs are inherited from metal-silicate melting and partial exchange before planetesimal accretion in a light isotope-enriched gas. Finally, the $\Delta^{74/70}\text{Ge}_{\text{metal}}$-$\Delta^{17}\text{O}_{\text{silicate}}$ correlation between the IIE iron meteorites and OCs, provides new evidence for the existence of a highly reduced HH group.

1. Introduction

Ordinary chondrites (OCs) are undifferentiated meteorites that formed in the first 1 to 3 Myr after the formation of refractory Calcium-Aluminium Inclusions (CAIs) (Kleine et al., 2008; Archer et al., 2019; Hellmann et al., 2019), as well as during a short accretion time interval of less than 0.5 Myr (Monnereau et al., 2013). OCs also postdate magmatic iron meteorite formation. Several studies argue for their formation via the disruption-re-accretion of early-differentiated planetesimals (Elkins-Tanton et al., 2011; Lichtenberg et al., 2018). As a result, OCs should represent a record between the early formed planetesimals and terrestrial planets as we know them today.

Ordinary chondrites are composed of an assemblage of chondrules and matrix, which contain silicates, FeNi alloys, sulfides, and rare CAIs. The OC sequence has been classified into three groups, i.e., H (High metallic Fe content), L (low metallic Fe content) and LL (low total Fe and low metallic Fe content), according to the oxidized iron content relative to the metal iron (Urey and Craig, 1953), oxygen isotopic composition (Clayton et al., 1991), bulk volatile elemental concentration, including water, and chondrule size (Wasson, 1972; Chou and Cohen, 1973; Chou et al., 1973; Kallemeyn et al., 1989; Wasson, 2000; Rubin, 2005). Each chondrite group is thought to originate from at least one parent body and are subdivided into petrologic types, from the least metamorphosed and unequilibrated (group 3) to the most metamorphosed
and highly equilibrated (group 6; Van Schmus and Wood, 1967). Studies on OC ages, cooling times, and petrographic types led to the creation of the "onion-shell" model in which thermal metamorphism increases with planetesimal depth (Göpel et al., 1994). With increasing petrographic type, the temperature (Tait et al., 2014) and oxidation state increase (Chou and Cohen, 1973; Chou et al., 1973; Rubin, 2005), metallic iron content decreases, and metal becomes more equilibrated with silicate through oxidation reactions (McSween and Labotka, 1993). Moreover, as thermal metamorphism increases, the metal grain-size, as well as the grain circularity, increase, which define additional parameters for characterizing the petrologic types (Guignard and Toplis, 2015).

The percentage of fayalite in olivine (%Fa) is commonly used as a redox proxy (Chou et al., 1973; Rubin, 2005). Clayton et al. (1991) first showed that there was an increase in the oxygen isotopic composition from the H to LL groups, which led McSween and Labotka (1993) and Rubin (2005) to demonstrate that there is a positive correlation between %Fa and the oxygen isotopic composition throughout the OC sequence. This implies that variations in δ18O and Δ17O are related to the oxidation state. However, the origin of oxidation processes remains highly controversial. Previous studies have argued that this evolution reflects primitive conditions in the solar nebula recorded by planetesimals through progressive accretion of high Δ17O nebula water hosted in phyllosilicates (i.e. Rubin, 2005). Thermal metamorphism, induced by heat release from the decay of short-lived radioactive elements (26Al), liberates heavy-O water, which produces both an increase in the oxidation state and oxygen isotopic composition. On the other hand, the oxidation state can also be explained by an increase in the proportion of type II chondrules (oxidized) and the decrease of type I chondrules (reduced) from H to LL. Type I is reduced, O16-rich, and contain metallic Fe, whereas type II is oxidized, 16O-depleted, and Fe is present in an oxidized form (Zanda et al., 2006). Chou and Cohen (1973) and Chou et al. (1973) have shown that OCs groups are also resolvable using the siderophile element concentration,
e.g., the germanium content in the metal increases from H to LL with the Ni content, as a proxy of the redox state. They proposed that a significant proportion of Ge condenses in the silicate at its formation, such that Ge is redistributed between the metal and silicate through oxidation-reduction during metamorphic processes on their parent bodies.

Germanium is a moderately siderophile and volatile element ($T_{50\%\text{ condensation}} = 883K$ at $\text{pH}_{2}\text{O}/\text{pH}_2 = 4.2 \times 10^{-4} \text{atm}$; Lodders et al., 2009), commonly used with Ir, Au, and Ni to constrain the formation of iron meteorites and metal phases in chondrites (Wasson, 1974). Experimental studies have shown that Ge partitioning between metal and silicate phases is strongly redox-dependent, with $D(\text{Ge})$ metal-silicate values greater than $10^4$ under highly reduced conditions ($-2.7 \log f_O^2$ relative to the IW buffer; Kegler and Holzheid, 2011). Germanium is also a chalcophile element but the amount of Ge in troilite (FeS) from the OCs represents only 3 to 5% (0.3 to 0.5 ppm) of the total amount of Ge in OCs (Chou and Cohen, 1973; Wai and Wasson, 1979), whereas it is strongly concentrated, i.e., up to several hundred ppm, in terrestrial low-temperature sulfides (Bernstein, 1985).

Germanium has five stable isotopes, i.e., $^{70}\text{Ge}$, $^{72}\text{Ge}$, $^{73}\text{Ge}$, $^{74}\text{Ge}$, and $^{76}\text{Ge}$, with the Ge isotopic composition given as a delta notation: $\delta^{74/70}\text{Ge} = (\frac{^{74}\text{Ge}^{70}\text{Ge}_{\text{Sample}}}{^{74}\text{Ge}^{70}\text{Ge}_{\text{Standard}}} - 1) \times 1000$. To date, previous studies have analyzed the germanium isotopic composition of differentiated bodies that experienced metal-silicate segregation to investigate the conditions and processes of planetary differentiation. Magmatic iron meteorites that represent planetesimal cores display positive and heavier Ge isotopic compositions than the Earth’s mantle and crust (Luais, 2007; Luais, 2012). The Ge isotopic composition of the Earth’s silicate reservoir is, thus far, homogeneous and has not been significantly modified by complex crust–mantle exchange via fluids in subduction zones (El Korh et al., 2017). The volatile behavior of germanium can also be responsible for large Ge isotopic fractionations that occur in IIE non-magmatic iron meteorites (Luais, 2007), which form due to metal segregation following impact-driven melting at the sub-
surface of their parental bodies (Wasson and Wang, 1986; Wasson, 2017). To date, the only constituents that display strong negative $\delta^{74/70}$Ge values are schreibersite (Fe-Ni-P) inclusions in iron meteorites (Rouxel and Luais, 2017) and low-temperature terrestrial Zn-sulfides (Belissont et al., 2014). Metal-silicate segregation, volatility, and phosphide/sulfide exsolution/crystallization are the main processes, currently known, that strongly fractionate germanium and its isotopes.

Germanium isotopes represent potential tools to investigate processes of formation and evolution of ordinary chondrites (1) during condensation of metal-silicate-sulfide from the solar nebula (Sears, 1978; Wai and Wasson, 1979), (2) by inheritance from previously destroyed planetesimals (Litchenberg et al., 2018), and (3) metal-silicate equilibrium fractionation on parent bodies following a major impact (Tomkins et al., 2013) or after metamorphism (Humayun and Campbell, 2002; Guignard and Toplis, 2015). In addition, ordinary chondrites provide a suitable environment to explore the relationships between Ge and O isotopes, as well as the redox conditions. The combined use of these data can establish constraints on the environment in which OCs form and the processes that lead to metal-silicate fractionation and equilibration. Germanium isotopic measurements on separated phases from unequilibrated and equilibrated OCs are used to evaluate the effect that metamorphism has on metal-silicate isotope fractionation. Clayton et al. (1991) first proposed the genetic association between ordinary chondrites and IIE iron meteorites based on their oxygen isotopic compositions in silicate clasts. The IIE–OCs relationship has been further constrained with Mo isotopic anomalies (Burkhardt et al., 2011; Budde et al., 2016). Here, based on the Ge isotopic composition in the metal phase, we examine a new genetic link between IIE iron meteorites and H OCs.

2. Samples and methods

2.1. Sample selection
To constrain the metamorphic and shock effects on Ge systematics, sixteen OCs (eight H, six L, and two LL) were selected to represent all petrologic types (Van Schumus and Wood, 1967) and various shock stages as defined by (Stöffler et al., 1991). In addition, two samples were specifically selected to assess the effect of shock: Rose City, an impact melt breccia, and Portales Valley, a low shock meteorite with large metal veins as an example of metal migration (Ruzicka et al., 2005). Table 1 summarizes the sample characteristics.

2.2. Analytical Procedures

2.2.1. Sample preparation

Approximately 500 mg of initial bulk meteorite material was cleaned with 0.2 M HNO\textsubscript{3} under cold ultrasonic agitation for 5 minutes. The acid was removed, the samples were rinsed with pure water under ultrasonic agitation and then rinsed with acetone to avoid metal oxidation. Afterward, 300 mg of the meteorite sample was gently powdered in an agate mortar for bulk elemental and isotopic measurements.

Metal and sulfide separation. The remaining 200 mg of clean sample was gently crushed in an agate mortar. FeNi metal was removed from the crushed sample with a hand-magnet. The magnetic fraction was crushed multiple times in an agate mortar, followed by several minutes of ultrasonic agitation in an acetone medium to remove all silicate particles remaining on the metal surface. A Frantz magnetic separator was used to separate the sulfides from the remaining weakly magnetic powder. Between 10 and 40 mg of metal and 10 to 20 mg of sulfide were separated using these procedures.

Silicate separation. The crushed non-magnetic sample fraction that mostly contained impure silicates was gently grounded in an agate mortar to a fine powder. To remove the non-magnetic sulfides and small metal particles stuck on the silicates, two successive cleanings with 6 M HCl
and 2 M HNO₃ were performed for 5 minutes under ultrasonic agitation at room temperature. As Ge is highly volatile when associated with HCl at a high temperature (T > 60°C), this protocol was first tested on San Carlos olivine samples to evaluate Ge loss. Uncleaned olivines have [Ge] = 0.89 ± 0.02 ppm and δ⁷⁴⁷⁰Ge = 0.70 ± 0.08‰ while clean olivines have [Ge] = 0.90 ± 0.02 ppm and δ⁷⁴⁷⁰Ge = 0.71 ± 0.09‰ (Table 2), which demonstrates that there was no Ge loss nor isotopic fractionation during cold HCl leaching. This step is important because metallic phases are at least one hundred times more concentrated in germanium than silicates or sulfides (Chou and Cohen, 1973). The presence of small metal particles on the silicate surface would have introduced a bias on the elemental and isotopic compositions of Ge in silicates.

Chondrule separation. Separation was only performed on Dhajala (H3.8). A total of 1 g of Dhajala was cleaned and gently crushed in an agate mortar to preserve the chondrules. Unbroken chondrules were separated by hand-picking and sieved into two fractions, i.e., 100–300 µm and > 300 µm, with a recovery of 9.1 and 130.5 mg, respectively.

All fractions were later checked and purified by hand-picking under a binocular microscope, successively cleaned with diluted HNO₃, distilled H₂O, and acetone under cold ultrasonic agitation, and dried under a laminar flow hood.

2.2.2. Sample digestion and chemical purification of germanium

The germanium chemistry analytical procedures for metal, sulfide, and silicate phases follow those described in Luais (2007, 2012, references therein) and Belissont et al. (2014). Samples were dissolved in Teflon beakers on a hotplate in concentrated HF-HNO₃ solutions for the silicate and sulfide fractions while metal phases were dissolved in an HNO₃-only solution. Bulk samples were digested under pressure (25 bar) using Bola Teflon bombs at 150°C (Luais et al., 2009) for one week in an oven to ensure the dissolution of all phases, including chromite. To
ensure that no Ge was lost at high temperatures under pressure, this protocol was tested using a BIR-1 geo-standard. We obtained a $\delta^{74/70}$Ge of $0.59 \pm 0.10\%$ for BIR-1 that was digested in the bomb while this was $\delta^{74/70}$Ge = $0.59 \pm 0.06\%$ for conventional hotplate BIR-1 dissolution. This demonstrates that there was no isotopic fractionation using this digestion method (Table 2).

Germanium purification from silicate and sulfide matrices was completed in two steps using ion exchange resins. First, anionic resin (AG1-X8 resin chloride form, 200 – 400 mesh, Biorad, Hercules, CA, U.S.A.) allowed us to eliminate most elements (i.e., the alkalis, Mn, Al, Co, Ni, Zn, Ga, and most Fe) using 1 M HF. Germanium (± Fe, Mg) was eluted with 0.2 M HNO$_3$. Germanium purification from Fe and the remaining matrix was performed solely using cation exchange resin (AG 50W-X8, Biorad, 159 Hercules, CA, U.S.A.) and 0.5 M HNO$_3$. Germanium isolation from the Fe-Ni matrix of the metal phase was performed only using the cation exchange resin step. These two procedures had a germanium yield of 100% within error (Luais, 2007; Luais, 2012). All acids used for digestion and germanium chemical separation were high-purity grade SeaStar® reagents.

2.2.3. Elemental measurements

Bulk powder samples and metal phase separates from H OCs, and metal from L samples were analyzed for major and trace elemental compositions including Ge at the SARM analytical facility (Service Analytique des Roches et des Minéraux, CRPG-Nancy) using a ThermoFischer ICP-MS X7. Errors associated with the Ge, Ni and Co measurement were < 5% for our sample concentrations. The separated mass of pure sulfides was low (i.e., approximately 10 – 20 mg) due to their low occurrence in the OCs (~5.7%; Lauretta et al., 1997). Similarly, only 10 – 40 mg of pure silicate separates were obtained. These available low mass fractions rule out the possibility of analyzing them using routine procedures at the SARM. These fractions were
analyzed for their Ge elemental compositions using a *ThermoFischer* ICP-MS *X-Series* at CRPG (Nancy, France) with a $2\sigma$ SD of 1.2%.

### 2.2.4 Ge isotopic measurements

Samples in a 0.5 M HNO$_3$ solution from the cation exchange step were analyzed for their Ge isotopic compositions using a Neptune*Plus* multiple-collector inductively coupled plasma mass spectrometer (MC-ICP-MS) at CRPG coupled with a hydride generator introduction system (HGIS). The HGIS method involves mixing the sample with a high-reducing solution of NaBH$_4$-NaOH. The reactions between these two solutions convert volatile aqueous species (GeOH$_4$) to gaseous (GeH$_4$) hydride species (Dedina and Tsalev, 1995; Abdul-Majeed and Zimmerman, 2012). As the reducing solution is in excess, the reaction yield is 100%. This technique removes isobaric interferences from elements that do not form hydrides, avoiding Zn, NiO, and FeO interferences, as well as matrix effects. The result is a gain in signal sensitivity of at least a factor of 30, required for routine isotopic measurements of 10 – 20 ng Ge (exceptionally down to 5 ng) to obtain high precision and a good external reproducibility of less than 0.1‰ ($2\sigma$ SD) (see Rouxel and Luais, 2017 for more details). Data were acquired in static mode using the $^{68}$Zn, $^{69}$Ga, $^{70}$Ge, $^{71}$Ga, $^{72}$Ge, $^{73}$Ge, and $^{74}$Ge isotopes that corresponded to the L3, L2, L1, C, H1, H2, and H3 Faraday cups, respectively. We report an intensity of 1.5 – 2.5 V on $^{74}$Ge for a 10 ppb Ge standard solution. Each measurement was composed of a 320 s data acquisition sequence distributed over 40 cycles. Procedural Ge blanks, including chemistry and MC-ICPMS, were better than 70 pg and negligible as this represents less than 0.1% of the 100 ng of Ge processed for each sample. Sample measurements were bracketed using the NIST3120a Ge standard solution. Results are reported in delta ($\delta$) notation using $^{x}$Ge/$^{70}$Ge ratios with respect to the mean ratios of the NIST 3120a standard analyzed just before and after the sample (Luais, 2012):
\[
\delta^{74/70}_{\text{Ge}_{NIST3120a}} = \delta \frac{x \text{Ge}}{\text{Ge}} = \left( \frac{x \text{Ge}/70 \text{Ge}}{(x \text{Ge}/70 \text{Ge})_{\text{NIST3120a}}} - 1 \right) \times 1000. 
\] (1)

JMC and Aldrich germanium standard solutions were analyzed during the course of the study, yielding \(\delta^{74/70}_{\text{Ge}} = -0.32 \pm 0.11\‰\) (n = 95) and \(\delta^{74/70}_{\text{Ge}} = -1.97 \pm 0.12\‰\) (n = 99) (2\(\sigma\) SD), respectively. The BIR-1 basalt geostandard yielded \(\delta^{74/70}_{\text{Ge}} = 0.63 \pm 0.17\‰\) (n = 10) and the in-house IAB Magura iron meteorite (Luais, 2007) had a \(\delta^{74/70}_{\text{Ge}}\) value of 0.78 ±0.08‰ (n = 19). These values agree with the references given in Luais (2012) and (Escoube et al., 2012).

3. Results

3.1. Elemental composition

The majority of the bulk H OCs displayed a rather uniform Ge concentration, ranging from 10.3 to 22.2 ppm (mean = 15.8 ± 0.8 ppm; see Table 3 and Fig. 1) with no variation among the petrologic types, which is in agreement with \([\text{Ge}]_{\text{mean}} = 12.7 \pm 1.5\ ppm\) reported in Fouché and Smale, (1967) and similar to the value reported for L OCs in Tandon and Wasson (1968). However, Allegan H5 ([Ge] = 22.2 ppm, [Ni] = 1.52 wt. %) and Guareña ([Ge] = 21.4 ppm, [Ni] = 1.59 wt. %) had slightly higher Ge and Ni contents (see Fig. A2 in the Appendix) compared to the OC mean ([Ge] = 15.8 ± 0.8, [Ni]$_{\text{mean}}$ = 1.28 ± 0.54 wt. %) and bulk data reported by (Wasson and Kallemeyn, 1988). For these samples, this may be due to a sampling bias from either a high taenite zone or, although the main mass may appear to be homogeneous, a small-scale heterogeneous repartition of metal that led to the sampling of a metal-rich zone.

Germanium concentrations in metal ranged from 50 to 83.5 ppm for H chondrites, with a mean of 64.6 ± 17.6 ppm (2\(\sigma\) SD), 104.3 to 157.8 ppm for L chondrites, with a mean of 135.9 ± 53.2 ppm (2\(\sigma\) SD), and 85 to 120 ppm for LL chondrites, with a mean of 102.5 ± 49.5 ppm. These values are within those reported by Chou et al. (1973) and Chou and Cohen (1973), who
proposed that a significant amount of Ge in the silicates was redistributed in the metal during metamorphism. Moreover, the increase in the oxidation state from H to L to LL should have decreased the amount of the metal phase, and, consequently, increased the concentration of Ge in the metal. Taken together, all the metal samples define a positive correlation among Ge, Ni, and Co, demonstrating that taenite (Ni-rich) is more concentrated in Ge than kamacite (Ni-poor) (see Figure A in the Appendix).

Germanium concentrations in silicates from the H group ranged from 0.2 (H4 Ste Marguerite) to 2.2 ppm (H3.4 Sharps) (see Table 3 and Fig. 1), identical to the order of magnitude for data reported in Chou et al. (1973). However, we observed a discrepancy for the Allegan Ge concentration between our data, i.e., 0.35 ppm compared with a value of 0.06 ppm reported in Chou et al. (1973). In the L group, Ge concentrations in silicates ranged from 0.05 ppm (L6 Alfanello) to 0.65 ppm (L4 Bjürbole) (Table 3 and Fig. 1). They are identical to the values reported by Chou and Cohen (1973).

The sulfides from H6 Kernouvé and H4 Ste Marguerite were strongly depleted in Ge, with concentrations of 1.75 and 1.95 ppb (see Table 3 and Fig. 1), respectively. This indicates the strong compatibility of Ge in the Fe-Ni metallic phase.

3.2. Germanium isotopic data

Germanium isotope measurements were performed on bulk samples, silicates, and sulfides from H group chondrites, whereas we only analyzed L samples for metal and silicate, and LL for metal (Table 3). All samples plot on the mass-dependent fractionation line (see Fig. B in the Appendix).

Bulk Ge isotopic compositions define a narrow range from $\delta^{74/70}\text{Ge} = -0.59 \pm 0.04\%$ to $-0.37 \pm 0.15\%$, with a mean of $\delta^{74/70}\text{Ge} = -0.54 \pm 0.09\%$. Only Dhajala (with the highest $\delta^{74/70}\text{Ge}$
value at \(-0.37\%\)) deviated by more than 0.1\% from the mean. Therefore, this sample was not included in the mean (see section 4.5 for explanation) (Fig. 2A).

The $\delta^{74/70}$Ge composition for metal from H group samples ranged from $-0.33 \pm 0.12\%$ (Dhajala) to $-0.57 \pm 0.12\%$ (metallic globules from Portales Valley), with a mean of $-0.51 \pm 0.09\%$ (N.B. we excluded the Dhajala value due to its unusual Ge isotopic composition, see section 4.5 for explanation). The $\delta^{74/70}$Ge$_{metal}$ compositions for L group samples showed small variations from $-0.35 \pm 0.09\%$ (Hedjaz) to $-0.28 \pm 0.01\%$ (Alfianello), with a mean of $\delta^{74/70}$Ge = $-0.31 \pm 0.06\%$. The LL group yielded a mean value of $\delta^{74/70}$Ge = $-0.26 \pm 0.09\%$.

Silicates displayed lighter Ge isotopic compositions than bulk samples, with a large range in $\delta^{74/70}$Ge from $-0.92\%$ for Allegan (H5) to $-0.50 \pm 0.11\%$ for Dhajala (H3), with a mean of $-0.67 \pm 0.33\%$ for the H groups (Fig. 2A) and from $-1.25$ (Alfianello, L6) to $-0.50 \pm 0.09$ (Björbole, L4) with a mean of $-0.87 \pm 0.18$ for the L group (Fig. 3A).

The sulfides were analyzed in two samples, without replicates due to their low abundance in OCs and Ge concentration ($\approx 2$ ppb). They were characterized by the lightest isotopic values at $\delta^{74/70}$Ge = $-1.60\%$ (Ste Marguerite, H4) and $\delta^{74/70}$Ge = $-1.56\%$ (Kernouvé, H6). We attributed the highest error to these samples based on the standards at $\pm 0.20\%$ (2σ SD) (Fig. 3).

The two chondrule fractions from Dhajala had similar isotopic compositions of $\delta^{74/70}$Ge = $-0.28 \pm 0.12\%$ for $< 300 \mu m$ chondrules and $\delta^{74/70}$Ge = $-0.38 \pm 0.11\%$ for $> 300 \mu m$ chondrules (Table 3).

3.3. Metal phases: main contributors to Ge isotopic compositions in bulk OCs

The metal phase is at least 60 times more concentrated in Ge than the silicates (Chou et al., 1973) and four orders of magnitudes higher than sulfide (Fig. 3A). The germanium isotope
compositions of bulk H chondrites (mean $\delta^{\text{74/70}}_{\text{Gebulk}} = -0.54 \pm 0.09\%$) were within error of the metal (mean $\delta^{\text{74/70}}_{\text{Gemetal}} = -0.51 \pm 0.09\%$), except for Sharps, and higher than the silicate (mean $\delta^{\text{74/70}}_{\text{Gesilicate}} = -0.67 \pm 0.33\%$). In a $\delta^{\text{74/70}}_{\text{Ge}}$ versus Log(1/Ge) diagram (Fig. 3B), the mixing line defined by the bulk sample, metal, silicate, and sulfide phases (mean $\delta^{\text{74/70}}_{\text{Ge sulfide}} = -1.56 \pm 0.02\%$), at all metamorphic grades, indicates that the $\delta^{\text{74/70}}_{\text{Ge}}$ measurements of separate phases are consistent with the $\delta^{\text{74/70}}_{\text{Ge}}$ measurements of the bulk chondrite. The Ge isotopic fractionation, i.e., $\Delta^{\text{74/70}}_{\text{Ge bulk-metal}} = \delta^{\text{74/70}}_{\text{Gebulk}} - \delta^{\text{74/70}}_{\text{Gemetal}} = -0.03 \pm 0.21\%$, for the H OC is near zero, compared with $\Delta^{\text{74/70}}_{\text{Ge bulk-silicate}} = 0.16 \pm 0.33\%$, $\Delta^{\text{74/70}}_{\text{Ge bulk-sulfide}} = 1.06 \pm 0.16\%$, and $\Delta^{\text{74/70}}_{\text{Gemetal-silicate}} = 0.22 \pm 0.36\%$ (Table 4; Fig. 4A,B). This indicates that the metal phase was the main contributor to the bulk Ge isotopic budget and that the $\delta^{\text{74/70}}_{\text{Gemetal}}$ values are a proxy for the $\delta^{\text{74/70}}_{\text{Gebulk}}$ values.

In addition, we did not observe any correlation between the degree of shock and germanium isotopic composition. For example, Allegan (H5, S1) and Roses city (H5, S6) had identical $\delta^{\text{74/70}}_{\text{Gebulk}}$ isotopic compositions (Table 3).

4. Discussion

4.1. Metal-silicate-sulfide fractionation during condensation

During the cooling of a primordial nebular gas, metal (FeNi) and forsterite (Mg$_2$SiO$_4$) condense at similar temperatures, i.e., $T_{50\% \text{ condensation}} = 1370K$ at $10^{-4}$ atm (Yoneda and Grossman, 1995). The most abundant gaseous molecular species of germanium are GeS, GeBr$_2$, and GeO (Sears, 1978). As GeBr$_2$ is a highly volatile species of Ge, which condenses at much lower temperatures than GeS and GeO (Sears, 1978; Wai and Wasson, 1979; Wood et al., 2019), here, we only discuss GeS and GeO condensation. Condensation of GeS and GeO via reactions with existing silicates and FeNi alloys occur at $T_{50\% \text{ condensation}} = 883K$ (ideal Ge
solution; Lodders et al., 2009) or $T_{50\% \text{ condensation}} = 825–830K$ (no ideal solution; Wai and Wasson, 1977; Wood et al., 2019) and $p_{\text{H}_2\text{O}}/p_{\text{H}_2} = 4.2 \times 10^{-4}$. Sears (1978) showed that Ge is siderophile during high-temperature condensation and becomes lithophile at low temperatures. Thus, germanium condenses first in the metal based on the following reaction (Sears, 1978; Wai and Wasson 1979):

$$\text{GeS}_\text{g} + \text{H}_2 = \text{Ge}_{\text{in metal}} + \text{H}_2\text{S}_\text{g}. \quad (2)$$

Then, at lower temperatures, the remaining Ge is incorporated into the silicate phase according to the following reaction (Sears, 1978; Wai and Wasson, 1977):

$$n\text{GeO}_\text{g} + 2\text{Mg}_2\text{SiO}_4\text{(s)} + \text{H}_2\text{O}_\text{g} = 2\text{MgSiO}_3\text{(s)} + \text{Mg}_2(1-n)\text{SiO}_4\text{(s)} + \text{H}_2. \quad (3)$$

On the other hand, sulfide condenses at lower temperatures than metal and silicate ($T_{50\% \text{ condensation}} = 691K$) via reactions between $\text{H}_2\text{S}_\text{gas}$ and solid metal based on the following reaction (Lauretta et al., 1997, and references therein):

$$\text{H}_2\text{S}_\text{g} + \text{Fe}_\text{g} = \text{FeS}_\text{s} + \text{H}_2\text{(g)}. \quad (4)$$

The extent and direction of isotopic fractionation in the solid condensate with respect to the vapor source is difficult to evaluate and depends on whether the system maintains equilibrium, is kinetically controlled, or is in a closed or open system. High-temperature equilibrium processes do not induce significant isotopic fractionation, whereas kinetic condensation can result in isotopic fractionation based on the relative cooling rate of the gas and partial pressure of the relevant elements in the gas (Richter, 2004). Calculations from Richter (2004) demonstrate that condensation in an environment, in which the temperature drop is relatively fast compared to the condensation time, results in $P_i < < P_{i,sat}$ (where $P_i$ is the partial pressure of a volatile element and $P_{i,sat}$ is its saturation vapor pressure). In this case, kinetic condensation produces a large isotopic fractionation between the source and condensing phase, where the first condensate is enriched in light isotopes. In the condensation sequence, the metal phase that condenses first has the
lightest Ge isotopic composition compared with that of the silicate and sulfide. In contrast, we observed higher $\delta^{74/70}_{\text{Ge, metal}}$ values for the metal than for the silicate and the lowest $\delta^{74/70}_{\text{Ge, sulfide}}$ values in the H chondrites (Fig. 3A).

We can draw the same conclusion from the sulfide data. Sulfides in equilibrated OCs are depleted in germanium, with a Ge content as low as 2 ppb and light isotopic composition of $\delta^{74/70}_{\text{Ge}} = -1.56\%$ (Fig. 3A). As the first Ge condensates are incorporated into the Fe-Ni alloys, the late sulfide condensate result in (i) a low Ge content (as observed here), despite the fact that Ge is a chalcophile element (Chou et al., 1973), and (ii) a heavier Ge isotopic composition for the sulfides as compared with the metal (Davis and Richter, 2014). The lighter $\delta^{74/70}_{\text{Ge, sulfide}}$ values compared with the $\delta^{74/70}_{\text{Ge, metal}}$ values, however, rules out condensation from a protoplanetary disk as a viable process for the origin of Ge isotopic fractionation in metal-silicates and sulfides in ordinary chondrites.

4.2. Evidence for metal-silicate-sulfide partial equilibration before OC accretion

Germanium isotopic compositions in metal, silicate, and sulfide phases in all samples of the H chondrites display the following relationship: $\delta^{74/70}_{\text{Ge, sulfide}} \ll \delta^{74/70}_{\text{Ge, silicate}} < \delta^{74/70}_{\text{Ge, metal}}$ (Fig. 3A). This relationship holds even for the most reduced type 3 samples, which are the least affected by thermal metamorphism at T<675°C (Tait et al., 2014), which is too low to melt metal or silicate. This observation suggests that the positive $\Delta^{74/70}_{\text{Ge, metal-silicate}} = +0.22 \pm 0.36\%$ (varying from +0.17 to +0.33\% in H3) values is due to metal-silicate exchange that occurred before the accretion of H ordinary chondrite parent body. This positive $\Delta^{74/70}_{\text{Ge, metal-silicate}}$ also occurs between the Earth silicate reservoir ($\delta^{74/70}_{\text{Ge}} = +0.56 \pm 0.16\%$) and magmatic Fe-meteorites ($\delta^{74/70}_{\text{Ge}} = +1.41 \pm 0.22\%$) (Luais, 2012), which may represent the core of early formed planetesimals. Therefore, higher $\delta^{74/70}_{\text{Ge}}$ values in metal compared to silicate, in both
undifferentiated and differentiated bodies, appears to be the result of metal-silicate equilibrium. The larger $\Delta^{74/70}$Ge$_{\text{metal-silicate}}$ of $+0.85\%$ in differentiated bodies reservoirs (e.g. the silicate Earth and Fe- meteorites, Luais, 2012) compared to $\Delta^{74/70}$Ge$_{\text{metal-silicate}}$ in OCs can reflect additional isotopic fractionation during core-mantle differentiation.

The addition of sulfur in the Fe-Ni system strongly modifies the Ge incorporation into the metal phase. Chabot et al. (2003) and Chabot and Jones (2003) demonstrated that, during fractional crystallization at FeNi-S metallic liquid equilibrium, the $D(\text{Ge})$$_{\text{solid metal/liquid metal}}$ value increases by several orders of magnitude as S increases in the liquid. Similar low Ge contents have also been observed in S-rich melts during Fe-Ni-S partial melting experiments (Rushmer et al., 2005). Low Ge contents (2 ppb) in sulfide from type 4 and 6 OCs (Fig. 3A) likely result from late-stage fractional crystallization of Fe-Ni-S melts. The strongly negative $\delta^{74/70}$Ge$_{\text{sulfide}}$ value of $-1.56\%$ appears to be a feature common to sulfides formed by fractional crystallization, such as terrestrial sulfides (Belissont et al., 2014). As we have no isotopic values for the H3 type, it is unclear if sulfide crystallization occurs during nebular processes or on the parent body. However, the discrepancy between temperatures associated with the H4 petrologic type (676–865°C; Tait et al., 2014) and FeNi-S fusion temperature (963°C; Mare et al., 2014) suggest that negative Ge isotopic composition of sulfides could have been inherited from processes that occur before parent body accretion, as it has been proposed for metal and silicate phases.

Equilibrium among metal, silicate, and sulfide in ordinary chondrites must then reflect processes before the accretion, such as (1) partial re-equilibration of material produced by disruption of partially differentiated planetesimals, or (2) a heating event in the solar nebula. Bischoff et al. (2018) reported that, in more than 2280 polished OC thin sections, there was a high abundance of brecciated rocks and occurrences of xenolithic fragments (e.g., CI, CM, and Ureilite) in type 3 OCs. This supports the hypothesis that OCs can represent aggregates of early formed achondritic, chondritic, and chondrule materials. Lichtenberg et al. (2018) argued that
OC parent bodies can form from early impact on a preheated (i.e., $^{26}\text{Al}$ decay) and partially differentiated planetesimal (Bryson et al. 2019). Thus, the negative $\delta^{74/70}\text{Ge}$ values of the bulk sample and all separated OC phases, compared with positive values for magmatic iron meteorites and Earth silicate reservoirs, imply a light isotope-enriched environment for OC formation, i.e., an impact plume environment. Evaporation processes that occur during an impact would likely overprint any $\Delta^{74/70}\text{Ge}_{\text{metal-silicate}}$ signature from a previous generation of planetesimals. Metal, metal-silicate, and sulfide would partially re-equilibrate after this event during partial melting in a gas enriched in light Ge isotopes. This model of metal melting before OC accretion agrees with a recent study of metal formation in L OCs (Okabayashi et al., 2019).

4.3. Post accretional processes

4.3.1. No shock effect on $\delta^{74/70}\text{Ge}_{\text{metal}}$ and $\delta^{74/70}\text{Ge}_{\text{bulk}}$ values

As the germanium is a moderately volatile element, evaporation should result in a significant enrichment in heavy isotopes in highly shocked samples. The H chondrite Roses City, an impact melt breccia (S6), shows no deviation in $\delta^{74/70}\text{Ge}_{\text{metal}}$ from the mean values for H group while Dhajala (S1) is enriched in heavy isotopes compared with other H chondrites. Similarly, for L group samples, there is no $\delta^{74/70}\text{Ge}_{\text{metal}}$ difference between highly shocked (Alfianello and Vouillé, both S5) and minimally shocked samples (Björbole, S1). The lack of a shock effect on the $\delta^{74/70}\text{Ge}$ values can be explained by the deep burial of material, which should inhibit light isotope evaporation due to the pressure (e.g., Roses City, H5), or an increase in temperature during a low energetic impact, which may be not sufficient to efficiently evaporate Ge.

4.3.2. Effects of metamorphism
Unequilibrated OCs (type 3) are primitive and may correspond to the outer layer of a layered planetesimal (Trieloff et al., 2003) that experienced low-temperature metamorphic processes (< 675°C; Tait et al., 2014). They represent the most reduced type of each OC group (Rubin, 2005). This implies that metal-silicate exchange is limited compared to other petrologic types and that younger processes may not have overprinted isotopic fractionation, which had been established before planetesimal accretion (Van Schmus and Wood, 1967).

From the H3 to H5 petrologic types, certain variations in the $\Delta^{74/70}$Ge metal-silicate values (Fig. 4B) indicate Ge isotopic exchange between metal and silicate during thermal metamorphism. However, the lack of variation in the $\delta^{74/70}$Ge$_{metal}$ values suggests that metamorphism had no measurable effect on the Ge isotopic composition of the metal phase. Thus, variations in the $\Delta^{74/70}$Ge$_{metal-silicate}$ values (Fig. 4B) indicate an increased sensitivity to $\delta^{74/70}$Ge$_{silicate}$ during metamorphic processes due to the lower Ge content in the silicate compared to the metal.

Melting points for the FeNi-S assemblage and silicate are 963°C and 1050°C, respectively, at 0.3 kBar (Mare et al., 2014 and references therein). Types 3 to 5 OCs experienced different peak metamorphism temperatures from $T < 675°C$ for H3, 525–727°C for H4, and 676°C < $T < 865°C$ for H5 (Monnereau et al., 2013; Tait et al., 2014). These temperature ranges are likely too low to melt FeNi-S alloy and cause metal-silicate exchange. In contrast, type 6 OCs reached temperatures between 866°C and 1000°C (Monnereau et al., 2013; Tait et al., 2014). FeNi-S melting and efficient Ge silicate-metal exchange resulted in an increase in the Ge content for metals from the H3–H5 types ([Ge]$_{mean}$ = 61.9 ppm), excluding Allegen (see section 3.1), to the H6 type ([Ge]$_{mean}$ = 66.0 ppm), which agrees with Ge diffusion from silicate to metal during metamorphism (Chou et al., 1973). Similarly, Ge isotopic exchange between metal and silicate in H6 resulted in a heavier silicate isotopic composition, which approached the isotopic composition of the metal. This led to a similar Ge isotopic composition for the silicate, metal, and bulk in the H6 chondrites within the 2σ error range (Fig. 4A,B).
The global scheme for L chondrites is not as clear as it is for H chondrites. Except for the Bjürbole sample (L4 at −0.5‰), silicates from the L6 chondrites have negative and heterogeneous $\delta^{74/70\text{Ge}}$ values that are distinct from the metal phase values (Figure 3A), which suggest that there was no isotopic re-equilibration, such as that observed in H6 chondrites. The “onion shell” model is less constrained for L chondrites. Göpel et al. (1994) proposed that the L parent body formed from thermal gradient disruption due to impact processes and reassembly. If this had occurred before metal-silicate equilibration, this could have prevented the efficient Ge exchange between the metal and silicate. Finally, the isotopic uniformity of the metal phases from H3 to H6 and L3 to L6 (Figure 2B) are consistent with closed-system metamorphism within each parent body.

4.4. Links between the redox process and $\delta^{74/70\text{Ge}}$ variations throughout the H–L–LL sequence

The main characteristic of ordinary chondrites is their increase in oxidation state through the H, L, and LL sequences, which is inferred from the decrease in metal content and increase in the percentage of fayalite in olivine (Chou et al., 1973). The similar increase in the $\delta^{18\text{O}}$ composition and $\Delta^{17\text{O}}$ anomalies from the H to LL OCs (Clayton et al. 1991) led Rubin (2005) to identify a correlation between the % Fa in olivine and the $\delta^{18\text{O}}$ and $\Delta^{17\text{O}}$ values.

The Ge concentration and isotopic composition in metal increases from the H to L to LL groups (Fig. 2B), which discounts the occurrence of any Ge loss and evaporative processes during OC formation. A well-defined positive correlation was identified among the %Fa, Ge concentrations, and $\delta^{74/70\text{Ge}}$ in metal for all petrographic types and groups (Fig. 5A and B). At first glance, this suggests that the metal becomes enriched in Ge and the heavy isotopes of Ge with increasing oxidation. However, as the D(Ge) value decreases with an increasing $fO_2$ (Kegler and Holzheid, 2011), we can expect the metal in the L chondrites to become depleted in Ge.
compared to the metal in H chondrites, whereas we observe the opposite (Fig. 5). Therefore, oxidation is not the main process that causes an increase in Ge content from H to LL. In the Urey and Craig-type diagram, \( \text{Fe}_{\text{metal}} + \text{sulfide} \) vs FeO (Krot et al., 2014), variations in reduced Fe and FeO do not follow the reduction-oxidation trend, suggesting that the system has lost some of its iron metal. In this case, the increase in the Ge content for the metal from H to LL, with a broadly constant Ge bulk concentration (Wasson and Kallemeyn, 1988), can be explained by the incorporation of Ge into the remaining metal. However, this process cannot explain the increase in the \( \delta^{74/70}\text{Ge} \) throughout the OC sequence. In conclusion, the increases in Ge, \( \delta^{74/70}\text{Ge} \), and %Fa are not related to oxidation conditions in the environment of OC formation, \( \text{i.e.} \), the \( f_\text{O}_2 \) conditions. One explanation may be the involvement of an oxidizing component.

The \( \delta^{74/70}\text{Ge} \) values and \( \delta^{18}\text{O} \) and \( \Delta^{17}\text{O} \) are also characterized by a positive correlation (Fig. 6) in which these three groups are distinguishable outside of the analytical errors, with the following values for H (mean \( \Delta^{17}\text{O} = 0.71 \pm 0.12\%e \) and \( \delta^{74/70}\text{Ge} = -0.51 \pm 0.09\%e \)), L (mean \( \Delta^{17}\text{O} = 1.02 \pm 0.15\%e \) and \( \delta^{74/70}\text{Ge} = -0.31 \pm 0.06\%e \)), and LL chondrites (mean \( \Delta^{17}\text{O} = 1.18 \pm 0.12\%e \) and \( \delta^{74/70}\text{Ge} = -0.26 \pm 0.09\%e \)). As metamorphism does not modify \( \delta^{74/70}\text{Ge} \) in metal and bulk chondrite, we suggest that Ge isotopic variations among the three groups are the result of accretionary processes. Hence, correlations among [Ge]-\( \delta^{74/70}\text{Ge} \)-%Fa, \( \delta^{18}\text{O} \), and \( \Delta^{17}\text{O} \) reflect processes that originated in the accretionary disk, which were broadly linked to the redox state. Specifically, chondrule mixing of type I-reduced and type II metal-free chondrules (Clayton et al., 1991, Zanda et al., 2006), as well as the incorporation of an oxidizing component during OC accretion (Rubin, 2005) were the two main processes (see discussion below).

4.4.1. Chondrules as carriers of a high \( \delta^{74/70}\text{Ge} \) signature?
Clayton et al. (1991) demonstrated that, in certain ordinary chondrites, small-size chondrules (with diameter $d < 300 \, \mu m$) show enrichment in $^{16}O$ and depletion in $^{17}O$ compared with larger chondrules ($d > 300 \, \mu m$). This difference has been interpreted in terms of the interaction between $^{16}O$-poor melted chondrules and $^{16}O$-rich gas at their time of formation. Chondrules were then classified into type I chondrules (magnesian), which are small, reduced, metal-rich, and depleted in volatiles, and type II chondrules (ferroan), which are larger, oxidized, metal-free, and volatile-rich (Zanda, 2004; Zanda et al., 2006). Additionally, Zanda et al. (2006) showed that the proportions of the type I chondrule decrease through the OC sequence while type II increases (type I: 45%, 28%, and 22%; type II: 36%, 58%, and 61% in the H, L, and LL groups, respectively). Based on the sizes of type I and II chondrules, their volume in the H, L, and LL OCs, and the oxygen isotopic compositions of large chondrules (heaviest) compared to small chondrules, Clayton et al. (1991) argues that the increase in $\Delta^{17}O$ through the H-L-LL sequence can be explained by a “size-sorting process” of chondrules, whereas Zanda et al. (2006) argue for mixing between petrological components of the chondrites. Thus, considering the $\delta^{74/70}Ge$-$\delta^{18}O$-$\Delta^{17}O$ correlations (Fig. 6 and 7), we can assess if type II chondrules have a heavier $\delta^{74/70}Ge$ composition compared with type I chondrules. An increase in type II can elevate the $\delta^{74/70}Ge$ ratio in type II rich-ordinary chondrites.

We used two approaches to test the chondrule mixing hypothesis. First, we calculated the inferred $\delta^{74/70}Ge$ composition of type I and type II chondrules, as well as the matrix, to account for variations in $\delta^{74/70}Ge$ from H to LL due to mixing between these components (Fig. 7A). This mass balance calculation is based on Zanda et al. (2006), with type I and II chondrules, the matrix proportions in H, L, and LL, and their respective oxygen isotopic compositions (see the caption for Fig. 7 for more details and Table 2 in Zanda et al. 2006). Based on this modeling, we obtained the following theoretical compositions: $\delta^{74/70}Ge_{\text{type I}} = -0.81\%o$, $\delta^{74/70}Ge_{\text{type II}} = 2.15\%o$, and $\delta^{74/70}Ge_{\text{matrix}} = -7.35\%o$. 


For the second approach, we analyzed the Ge isotopic composition of type I and II chondrules. As it is not possible to distinguish between type I and II chondrules at the macroscopic scale, we used the same protocol as Clayton et al. (1991) on the same meteorite (Dhajala H3.8). We separated two chondrule fractions (i.e., \(d < 300\) and \(d > 300\) \(\mu m\)) based on the assumption that the smaller fraction should be mainly composed of type I chondrules while the larger fraction contains type II chondrules. Our results on the chondrule fractions show that the \(\delta^{74/70}\text{Ge}\) values of small \((\delta^{74/70}\text{Ge}_{d<300\mu m} = -0.28 \pm 0.12\%o)\) and large chondrules \((\delta^{74/70}\text{Ge}_{d>300\mu m} = -0.38 \pm 0.11\%o)\) (Table 3) are similar within error and are near the Dhajala bulk value \((\delta^{74/70}\text{Ge} = -0.37 \pm 0.15\%o)\). In the \(\delta^{74/70}\text{Ge}-\delta^{18}\text{O}\) diagram, the graphic, Dhajala bulk chondrite plots on the mixing line defined by type I and II Dhajala chondrules, indicating no analytical artefact. However, type I and II chondrule isotopic compositions do not plot on the \(\delta^{74/70}\text{Ge}-\delta^{18}\text{O}\) OC mixing line (Fig. 7B) and do not agree with the calculated values for type I and II chondrules. This precludes any mixing process between type I and II chondrules that can explain the H, L, and LL correlation.

An in-situ ion probe study of oxygen isotopes in individual olivine and pyroxene phases of OC chondrules yielded a more complex picture. Variable \(\delta^{18}\text{O}\) but similar \(\Delta^{17}\text{O}\) values have been found in type I and II chondrules from H chondrites, as well as the Semarkona (LL3) chondrite (Kita et al., 2010). In Semarkona, type I and II chondrules both have similar \(\Delta^{17}\text{O}\) values (+0.5–0.7\%o) that are lower than bulk LL chondrites (+1.2\%o), indicating that chondrule phenocrysts do not host the \(\Delta^{17}\text{O}\) anomalies. Thus, the issue is not the dichotomy between the O isotopic compositions of type I and II chondrules but the recognition of a high \(\Delta^{17}\text{O}\) anomaly. For example, a previous study has found a \(\Delta^{17}\text{O}\) anomaly in chondrule glass from the LL chondrite, which resulted from aqueous alteration with a high \(\Delta^{17}\text{O}\) ice component (Kita et al., 2010). Germanium isotopes and detailed studies of O anomalies in chondrules do not favor a mixing model that links phenocrysts in type I and II chondrules with the redox state but mixing with, or
incorporation of, a high $\Delta^{17}$O oxidized component in the chondrules or chondrite matrix is a viable model. The similar Ge isotopic composition between the type I and II chondrules suggests that evidence of an oxidizing component with high $\Delta^{17}$O must occur in the matrix of ordinary chondrites. A systematic investigation of the Ge isotopic compositions of chondrules and the matrix of the H, L, and LL types is necessary to understand similarities in their composition.

4.4.2. $\delta^{74/70}$Ge- $\Delta^{17}$O-%Fa correlations in OCs and accretion of hydrated phyllosilicate

The main issue with the correlation between Ge elemental/isotopic variations in ordinary chondrites and %Fa in olivine, as well as between $\delta^{18}$O and the $\Delta^{17}$O, is the origin of oxidation and how it relates to asteroidal processes and/or a nebular origin.

Clayton and Mayeda (1988) first identified the correlations between %Fa in olivine and $\delta^{18}$O and $\Delta^{17}$O in Urelites. They interpreted these correlations to be associated with oxidation processes in the nebula. Similarly, Rubin (2005) highlighted the $\Delta^{17}$O-%Fa positive correlation among OC groups, suggesting that the increase in the oxidation state through the OC sequence is related to an increase in the proportion of an accreted oxidizing agent with high $\delta^{17}$O, $\delta^{18}$O, and $\Delta^{17}$O signatures. Choi et al. (1998) identified such elevated oxygen isotopic signatures in magnetite from unequilibrated LL OCs, which show a large and unusual deviation from the Terrestrial Fractionation Line (TFL; $\Delta^{17}$O = 5‰) compared with OCs that have typical values ($\Delta^{17}$O ≈ 0 to 2‰). Choi et al. (1998) suggested that high $^{18}$O-nebular water ($\delta^{18}$O = 21‰ and $\Delta^{17}$O ≈ 7‰) accreted onto the OC parent body and induced low-temperature aqueous alteration. Such asteroidal alteration was responsible for the formation of magnetite via FeNi alloy oxidation (Wasson and Krot, 1994; Krot et al., 1997; Choi et al., 1998), the growth of fayalite rims on pyroxene grains via reactions between oxidized Fe and silica (Wasson and Krot, 1994), and the
presence of phyllosilicates in LLs and unequilibrated OCs (Hutchison et al., 1987; Choi et al., 1998).

Attributing the high Ge-$\delta^{74/70}$ compositions to the high $^{18}$O-nebular component (Figs. 5 and 6) and the main oxidizing agent in the low temperature ($T < 500K$) nebula to an $H_2O$-gaseous component (McSween and Labotka, 1993; Krot et al., 1997) is tempting. However, there is a large debate on the thermodynamic state of $H_2O$. McSween and Labotka (1993) and Krot et al. (1997) have proposed that ice condensate in the nebula accreted to OC parent bodies, and, then, later heated to form vapor during metamorphism and reacted to form phyllosilicates, such as described in the LL3 unequilibrated chondrites. Wasson and Krot (1994) and Wasson (2000) argued that the low temperatures, i.e., 160 K (or at 182 K; Lodders et al., 2009), required to condensate ice cannot be attained in the nebular locations that form OCs (within 2–3 AU). Instead, they have suggested that $H_2O$ derived from accreted hydrated silicates. Here, the key issue is how germanium can be incorporated into the low-temperature $H_2O$-oxidizing component. At germanium condensation temperature ($T_{50\%\text{ cond}}$) of 875K at $10^{-4}$ atm, Ge may have already condensed, and low-temperature ice condensates likely do not contain much Ge. However, hydrated silicates that form at the end of the condensation sequence but at higher temperatures than ice ($T_{50\%\text{ cond}} < 350°C$; Grossman and Larimer, 1974) can contain small amounts of Ge. Therefore, one possibility is that dust enrichment in the continuously accreting nebula can yield higher $T_{50\%\text{ cond}}$ condensation (Wasson, 2000; Davis and Richter, 2014). We propose that the last fraction of Ge, likely in the form of GeBr$_2$ (the second Ge most abundant species of Ge specie in the solar nebula, which is highly volatile; Sears, 1978), can condense in phyllosilicates and acquire the high $\delta^{74/70}$Ge values discussed in section 4.1

We have modelled the mixing between bulk H chondrite and the oxidizing phyllosilicate (Fig. 8) to (1) reproduce the OCs correlation in Figure 6 and, (2) determine what percentage of
accreted phyllosilicate is needed to change the composition from H to LL chondrites in terms of Ge content and $\delta^{74/70}$Ge in phyllosilicate. To constrain the model, we have considered the amount of accreted phyllosilicate, which cannot be higher than the amount of matrix (<15%, Krot et al., 2014), and the Ge content in terrestrial phyllosilicates. This value ranges between 1.1 – 8.5 ppm with means of 3.8 ± 3.2 ppm (2σ SD, calculated from Bernstein, 1985, and references therein). The model shows that at Ge content of 3.8 ± 3.2 ppm and ~10 to 15% of accreted phyllosilicate, a composition from $\delta^{74/70}$Ge = +6‰ to +3.2‰, respectively, is needed to change the isotopic composition from H to LL. By extrapolation of the $\delta^{74/70}$Ge-Δ$^{17}$O correlation shown in Figure 6, we find that the $\delta^{74/70}$Ge = 3.2 – 6‰ in phyllosilicates would correspond to a Δ$^{17}$O ≈ +8 to +15‰. The lower Δ$^{17}$O is close to the value of the oxidized nebular component incorporated in OC parent bodies reported by Choi et al. (1998) (Δ$^{17}$O ≈ +7‰), yet a better match is attained for a $\delta^{74/70}$Ge = +2.5‰ (Δ$^{17}$O = +7‰) in phyllosilicates. This value implies a Ge content of about 4.8 to 7 ppm for 15% and 10% of phyllosilicate, respectively. So, it seems possible that phyllosilicates could contain between 4 to 7 ppm of Ge with a composition of $\delta^{74/70}$Ge = +3 to +2.5‰. It can be noted that a Ge content in phyllosilicates of few ppm is likely from the thermodynamic calculation (Sears, 1978). They predict that a significant proportion of Ge (in the form of GeBr$_2$) would remain in the gas even after troilite condensation and that Ge would become lithophile at the end of the condensation sequence (Sears, 1978).

During planetesimal heating, the dehydration of accreted phyllosilicates triggers a redox reaction on the parent body. The release of $^{18}$O-rich and $\delta^{74/70}$Ge-high water (or vapor) oxidizes the Fe-Ni metal as FeO, forming magnetite, which then increases the %Fa in olivine. Germanium released in water is likely incorporated in the metal due to its siderophile behavior. However, the small amount of Ge does not contribute to the significant increase in Ge from the H to L to LL OCs. In contrast, the high $\delta^{74/70}$Ge composition of the released H$_2$O component that reacted with the metal should have significantly increased the $\delta^{74/70}$Ge$_{\text{metal}}$ proportional to the
amount of accreted oxidizing agent on each OC parent body. As such interaction may occur
during metamorphism, a correlation for the $\delta^{74/70}_{\text{Ge}_\text{metal}}$ composition is possible for types 3 to 6 in
a given OC group. However, we propose that the lack of a $\delta^{74/70}_{\text{Ge}_\text{metal}}$ correlation with the
petrologic type is imputable to the relatively high temperature of thermal metamorphism, which is
even recorded in the earlier stage of pervasive H$_2$O alteration via magnetite formation in the type
3 chondrite (T < 675°C) (Krot et al. 1997; Choi et al., 1998; Tait et al., 2014). The fractionation of
stable isotopes at equilibrium is a function of $1/T^2$, such that an increase in the metamorphic
temperature from types 3 to 6 (T < 675 to 975°C, Tait et al., 2014) does not induce significant
variations in the Ge isotopic fractionation between the H$_2$O-component and metal with the
petrologic type. Hence, the observed Ge isotopic fractionation among groups is primarily a
function of the amount of nebular hydrated phyllosilicate with a high $\delta^{74/70}_{\text{Ge}}$ signature accreted
onto the parent body. The heavy O isotopic composition of OCs that partly derive from
phyllosilicate destabilization can be considered, along with the previous H$_2$O gas or ice model for
the origin of accreted water.

4.5. Dhajala: A reduced L chondrite

The higher $\delta^{74/70}_{\text{Ge}_\text{metal}}$ composition of Dhajala (H3.8) compared with the other H chondrites,
including Sharps (H3.4), which has a similar petrologic type, requires discussion because
Dhajala does not fit the correlations shown in Figs. 5B and 6. Dhajala and Sharps have both
experienced a low degree of metamorphism. They lie within the oxygen isotope field as H
chondrites but Dhajala (H3.8) has oxygen isotopic anomalies that are slightly higher than Sharps
(H3.4) ($\Delta^{17}$O$_{\text{Dhajala}} = 0.698$ ‰ and $\Delta^{17}$O$_{\text{Sharps}} = 0.646$ ‰; Clayton et al., 1991). The high Fe$_{\text{metal}}$
(15.98 wt. %), and low FeO$_{\text{bulk}}$ (10.18 wt. % ) contents in Dhajala compared with Sharps (Fe$_{\text{metal}}$
= 12.02 wt. % and; FeO$_{\text{bulk}}$ = 13.64 wt. %) (Table 5) reflect more reduced conditions for Dhajala
than for Sharps. In contrast, the high percentage of fayalite and low carbon content in Dhajala
(\%\text{Fa} = 19.3; C = 0.04 \text{ wt. \%}) indicate more oxidized conditions than for Sharps (\%\text{Fa} = 17.5 and 
C = 0.95 \text{ wt. \%}) (Jarosewich, 1990). In addition, the Dhajala germanium isotopic
composition is closer to L type composition than H type (Figs. 2 and 6). Thus, we propose that
Dhajala accreted identical proportions of oxidizing component than L chondrites, which led to
olivine formation with high \%\text{Fa} (> 20) and an increase in the $\delta^{74/70}\text{Ge}_{\text{metal}}$ composition that
approaches the L Ge isotopic signature. Dhajala then experienced reduction by carbon
(McSween and Labotka, 1993) based on the following reaction (Warren and Huber, 2006):

$$\text{C} + \text{MgFeSiO}_4 = \text{MgSiO}_3 + \text{Fe} + \text{CO}.$$  (5)

In conclusion, the elevated $\delta^{74/70}\text{Ge}_{\text{metal}}$ composition in Dhajala with low \%\text{Fa} is the result of
high carbon activity in this meteorite, which did not decrease due to C dissolution in taenite
(Romig and Goldstein, 1978) by contrast with Sharps. This implies that Dhajala most likely
formed in the L region of the accretionary disk before reduction via carbon. This suggests that
the OC accretion environment is heterogeneous and $\delta^{74/70}\text{Ge}$ should be considered a tracer of
the primary locations that formed undifferentiated meteorites.

4.6. Relationship between the OC redox state and IIE iron meteorites

IIE non-magmatic iron meteorites formed by impact-induced melting and the destruction of a
chondritic parent body, followed by metal silicate re-accretion (Luais, 2007; Ruzicka, 2014;
Wasson and Wang 1986). This group can be divided into old (Miles, Weekeroo Station) and
young groups (Netschaëvo, Watson), which formed from 4.3 – 4.5 and 3.5 – 3.8 Ga, respectively
(Bogard et al., 2000). They have a complex mineralogy, mainly composed of metal and silicate
inclusions. These inclusions represent 5 to 10 vol. \% of the IIE meteorites and are composed of
silicates (feldspar, pyroxene, and olivine) phosphates, sulfides, and chromite (McDermott et al.,
Silicate inclusions can be classified into two categories, chondritic inclusions with relic chondrules and differentiated inclusions (McDermott et al., 2016).

The genetic links between the H OC and IIE iron meteorites were first identified based on the similar elemental and oxygen isotopic compositions of silicate inclusions from IIE iron and silicates from the H OC (Olsen and Jarosewich, 1971; Clayton and Mayeda, 1996). However, based on the (1) concentrations in the refractory and volatile siderophile elements and (2) low $\Delta^{17}$O values in silicate inclusions in the IIE Netschaevo meteorite, we can infer that the impact, which formed the IIE iron meteorites, occurred on an HH parent body that was more reduced than the H group (McDermott et al., 2016; Wasson, 2017).

To study the genetic link between the IIE irons and OCs, we only must consider meteorite samples from the IIE old group because they display the most pristine $\delta^{74/70}$Ge$_{\text{metal}}$ compositions compared with the young group, which show an increase in $\delta^{74/70}$Ge$_{\text{metal}}$ due to successive impacts (Luais, 2007). Figure 9 shows the positive correlation among the $\delta^{74/70}$Ge composition of IIE bulk metal (Luais, 2007; Luais, 2012), OC metal (data from this study), and $\Delta^{17}$O measured in silicate inclusions from IIE irons and bulk silicates from the H, L, and LL OCs (see Table 5 for oxygen references). IIE iron meteorite values for $\delta^{74/70}$Ge$_{\text{metal}}$ and $\Delta^{17}$O$_{\text{silicate inclusions}}$ are lower than those for the H chondrites. As the H chondrites are the less oxidized group, the lower $\delta^{74/70}$Ge$_{\text{metal}}$ and $\Delta^{17}$O$_{\text{silicate inclusion}}$ values (Fig. 9) of the IIE old groups compared to the H group, correspond to the isotopic composition of the most reduced HH parent body. Germanium isotopes demonstrate that there is a genetic link between the IIE iron meteorites and HH parent body.

5. Summary
We report the first data set of germanium isotopic compositions for bulk and separated phases from ordinary chondrites. The Germanium isotopic compositions were measured by an MC-ICP-MS coupled to a hydride generator system with a 2σ long term reproducibility of 0.1‰. We found that the Ge isotopic composition of ordinary chondrites is mainly driven by the metal phase, which is a good proxy for the OC bulk composition. The $\delta^{74/70}$Ge$_{metal}$ values significantly increased throughout the OC sequence, such that H, L, and LL groups are distinguishable at $-0.51 \pm 0.09$‰, $-0.31 \pm 0.09$‰, and $-0.26 \pm 0.09$‰, respectively. The implications of the germanium isotopic compositions for the understanding of ordinary chondrite compositions can be summarized as follows.

- The lack of significant $\delta^{74/70}$Ge$_{metal}$ variation between the unequilibrated and equilibrated OCs within each group implies that $\delta^{74/70}$Ge variation among groups is inherited from nebular processes rather than parent body processes.
- Metal and silicate are partially equilibrated in a closed system during metamorphism.
- The positive $\Delta^{74/70}$Ge$_{metal-silicate}$ signature is remarkably similar to previously reported Ge isotopic fractionations between magmatic iron meteorites and Earth silicates. However, the light $\delta^{74/70}$Ge composition of both metal and silicate suggests that metal, silicate, and sulfide have interacted before the accretion of ordinary chondrite parent bodies in a gas enriched in light Ge isotopes.
- The two well-defined positive correlations between $\delta^{74/70}$Ge$_{metal}$ and %Fa, as well as $\Delta^{17}$O, demonstrate that the Ge isotopic composition can be used as a tracer of oxidizing processes during accretion. The similar Ge isotopic compositions of small and large chondrules (from Dhajala) indicate that the size sorting chondrule mechanism cannot account for the $\delta^{74/70}$Ge$_{metal}$-%Fa and $\Delta^{17}$O correlations. Therefore, these correlations can be explained by an increase in the amount of accreted high $\Delta^{17}$O hydrated nebular phyllosilicates with high $\delta^{74/70}$Ge of about 2.5 to 3 ‰, throughout the ordinary chondrite
sequence. We can reconsider the origin of the heavy O isotopic signature in ordinary chondrites to derive from either H₂O (water) or hydrated silicate.

- The $\delta^{74/70}\text{Ge}_{\text{metal}}\Delta^{17}\text{O}_{\text{silicate}}$ correlation in the OCs provides evidence for the existence of an HH reduced parent body and its genetic link with the IIE iron meteorites.

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Research Data

Research Data associated with this article can be access at https://doi.org/10.17632/5rfdkx27w.1

References


Figure Captions

**Figure 1:** (A) Elemental compositions of the bulk meteorites (filled circles), metal (diamond), silicate (empty circle), and sulfide (square) from H (in blue), L (in red), and LL (in green) ordinary chondrites. Data shows large variation among all phases. The germanium concentration decreases from the metal (60 to 170 ppm as a function of the chondrite group) to bulk (≈10 to 20 ppm) to silicate (≈0.1 to 1 ppm) to sulfide (≈2 ppb) phases. There is no variation associated with the petrologic type. (B) A detailed view, in log scale, of the low Ge silicates and sulfide phases (dashed rectangle of A). Table 1 lists all abbreviations.

**Figure 2:** (A) A detailed view of the δ⁷⁴/⁷⁰Ge isotopic compositions among the bulk, metal, and silicate in the H ordinary chondrites. The meteorites are ordered from the least metamorphosed (Dhajala H3.4) to highest (Guaréña H6) types. Rose City and Portales Valley are plotted separately due to their specific textures (see text). The δ⁷⁴/⁷⁰Ge compositions become lighter from metal, to bulk sample, and, finally, to silicate. Silicates display larger δ⁷⁴/⁷⁰Ge variations than the metal and bulk samples. Isotopic fractionation between the phases are: Δ⁷⁴/⁷⁰Gebulk-metal = −0.03 ± 0.21‰, Δ⁷⁴/⁷⁰Gebulk-silicate = 0.16 ± 0.33‰, Δ⁷⁴/⁷⁰Gebulk-sulfide = 1.06 ± 0.16‰, and Δmetal-silicate⁷⁴/⁷⁰Ge = 0.22 ± 0.36‰. (B) A comparison of the δ⁷⁴/⁷⁰Ge metal values between the H (blue diamonds), L (red diamonds), and LL (green diamond) groups. All meteorites from the H and L and LL groups are plotted, with mean values of Hmean = −0.51 ± 0.09‰, Lmean = −0.31 ± 0.06‰, and LLmean = −0.26 ± 0.09‰. The value for Dhajala was not included in the mean value for H due to its off-trend isotopic composition (see section 4.5 for more details). Identical abbreviations as Fig. 1 and Table 1.

**Figure 3:** (A) δ⁷⁴/⁷⁰Ge versus log(Ge) concentration of the bulk and mineral phases of the H, L, and LL ordinary chondrites. (B) Negative correlation between δ⁷⁴/⁷⁰Ge and log(1/[Ge]) for the metal, silicate, bulk sample, and sulfide phases in the H chondrites (R² = 0.85).

**Figure 4:** Variations in the germanium isotopic composition and isotopic fractionation for H ordinary chondrites. (A) Correlation between Δ⁷⁴/⁷⁰Ge金属-硅酸盐 and Δ⁷⁴/⁷⁰Ge壳-硅酸盐 with a slope of 1.2 (R² = 0.9). Metal and bulk Ge isotopic compositions are similar within a 2σ SD. (B) Δ⁷⁴/⁷⁰Ge金属-硅酸盐 versus δ⁷⁴/⁷⁰Ge金属 shows no isotopic variation for metals among the petrologic groups. Values are from Table 4.

**Figure 5:** (A) The Ge elemental composition versus %Fa in olivine for the three OC groups. All data plot near the regression line (R² = 0.92). (B) δ⁷⁴/⁷⁰Ge of the metal phase versus %Fa in olivine for the H, L, and LL groups. Both the H and L groups plot near the regression line, where no clear correlation occurs within the groups. There
is no data for %Fa for Parnalle due to its large range in olivine composition (e.g., Rubin, 1990). Ge data for LL for Sako-Banja (SB), Cherokee Spring (CS), Olivenza, (Olz), Paragould (Prg), Dhurmsala (Drm), and Mangwendi (Mgw) are from Chou and Cohen (1973) and %Fa data are from Dodd (1968) and Dunn et al. (2010). See Table 5 for other references and %Fa.

Figure 6: δ74/70Ge of the metal phase versus Δ17O of the bulk sample for type 3 to 6 ordinary chondrites. Δ17O values are from Clayton et al. (1991). Rectangles in color represent fields of Δ17O variations (1σ SD) calculated from data reported in Clayton et al. (1991) and δ74/70Ge variations (2σ SD). Note that Dhajala value was not included in the H mean because due to its special isotopic composition (see section 4.5 for more details). The dotted line is a linear regression calculated for all individual values (R² = 0.676). %Fa, Δ17O data, and references are given in Table 5.

Figure 7: (A) Germanium isotopic compositions versus oxygen isotopic compositions for the two chondrule fractions from Dhajala (< 300 μm are filled blue diamonds and > 300 μm are empty blue diamond). The x-axis oxygen isotopic compositions for the < 300 μm and > 300 μm size fraction chondrules are the variations reported in Clayton et al. (1991). Blue, red, and green circles represent averages of the germanium and oxygen isotopic compositions for the H, L, and LL groups, respectively. Oxygen isotopic compositions and references are given in Table 5. (B) Theoretical Type I and II chondrules and matrix germanium isotopic composition based on chondrules and matrix mixing calculations. All data for the chondrules, matrix proportions (type I = 45%, 28 % and 22%; type II = 36%, 58%, and 61%; matrix = 12.4%, 10.2%, and 15.4%, for H, L, and LL, respectively), and their respective oxygen isotopic compositions are from Zanda et al. (2006). The calculated Ge isotopic compositions of δ74/70Ge_{type I} = –0.81‰, δ74/70Ge_{type II} = +2.15‰, and δ74/70Ge_{matrix} = –7.35‰ do not agree with the measurements of δ74/70Ge in chondrules displayed in Figure 7A.

Figure 8: Modelling of the Germanium content (ppm) in phyllosilicate required to change the germanium isotopic composition from an H (–0.51‰) to LL chondrite composition (–0.26‰) as a function of the accreted amount of phyllosilicates and their δ74/70Ge_{phyllosilicate} composition at 1‰, 2‰, 3‰, 4‰, 5‰, 10‰, 15‰, and 20‰. The blue dashed line corresponds to δ74/70Ge_{phyllosilicate} = +2.5‰, representing the isocurve that matches the Δ17O of ≈+7‰ reported in Choi et al. (1998) (see text). The horizontal dark blue line is the mean Ge content in terrestrial phyllosilicates (3.8 ppm; Bernstein, 1985, and references therein) and the dark blue zone is the 2σ SD of the mean. The highlighted area in black represents the likely composition of the phyllosilicates (see the text for more
The forbidden zone can be defined beyond the maximum percentage of matrix in ordinary chondrites (15%; Krot et al., 2014) and maximum Ge content in terrestrial phyllosilicates (up to 8.5 ppm; Bernstein, 1985).

**Figure 9:** $\Delta^{17}\text{O}$ of silicates (bulk ordinary chondrites and silicate inclusions of IIE iron meteorites) versus $\delta^{74/70}\text{Ge}$ of the metal phase in ordinary chondrites and IIE iron meteorites. The positive correlation between the IIE iron meteorites and LL ordinary chondrites demonstrates that the IIE parent body could have formed on a distinct and more reduced parent body than the H ordinary chondrite body, known as the HH parent body. As the $\Delta^{17}\text{O}$ of the NWA-118 sample is low compared to other LL chondrites, we also report the mean $\delta^{74/70}\text{Ge}$ versus $\Delta^{17}\text{O}$ for the LL chondrites (green square). References for the oxygen isotopic compositions are given in Table 5. The Ge isotopic composition of the IIE iron meteorites are from Luais (2007, 2012).
Appendix

**Figure A:** Germanium, nickel, and cobalt concentrations of the bulk and metal from H ordinary chondrites. (1) Co versus Ni: the Allegan (H5) metal phase exhibits high Ni and low Co contents that are characteristic of a taenite phase compared with kamacite. (2) Ge versus Ni: the high Ge contents of Allegan suggest a nugget effect that characterizes the Ni-rich, Co-poor taenite phase. Same legend for (1) and (2).

**Figure B:** $\delta^{74/70}\text{Ge}$ versus $\delta^{72/70}\text{Ge}$ of bulk, silicates, metal, and sulfide from the H group, and the metal phase from the L and LL groups of the ordinary chondrites. Error bars are $2\sigma$ SD. Except for sulfides, all standards and samples fall within error of the mass-dependent fractionation line, which is also the terrestrial mass fractionation line (TMFL) (Luais, 2007). Large errors for sulfide result from the small amount of measured Ge due to its low Ge concentration. Ge isotopic compositions become lighter from metal to bulk, silicate, and, then, to the sulfide phase. Identical trends as the $\delta^{74/70}\text{Ge}$ versus $\delta^{73/70}\text{Ge}$ diagram.
Figure 1, Florin et al., 2019
Figure 2, Florin et al., 2019
Figure 3, Florin et al., 2019
Figure 4, Florin et al., 2019
Figure 5, Florin et al., 2019
Figure 6, Florin et al., 2019
Figure 7, Florin et al., 2019
Figure 8, Florin et al., 2019
Figure 9, Florin et al., 2019
Table 1. The H, L, and LL ordinary chondrites analyzed in this study, with their petrologic types and shock stages. All meteorites are falls.

<table>
<thead>
<tr>
<th>Samples</th>
<th>Sample N° (Sources)</th>
<th>Petrologic types</th>
<th>Shock stage and references</th>
</tr>
</thead>
<tbody>
<tr>
<td>Sharps (Shs)</td>
<td>USNM 640 (SI)</td>
<td>H3.4</td>
<td>S3 (1)</td>
</tr>
<tr>
<td>Dhajala (Dhj)</td>
<td>USNM 5832 (SI)</td>
<td>H3.8</td>
<td>S1 (1)</td>
</tr>
<tr>
<td>St Marguerite (StM)</td>
<td>MNHN 3290 (MNHN-Paris)</td>
<td>H4</td>
<td>S2 (2)</td>
</tr>
<tr>
<td>Allegan (Alg)</td>
<td>USNM 215 (SI)</td>
<td>H5</td>
<td>S1 (2)</td>
</tr>
<tr>
<td>Rose city (RC)</td>
<td>USNM 6771 (SI)</td>
<td>H5 IB</td>
<td>S6 (1)</td>
</tr>
<tr>
<td>Guarena (Gn)</td>
<td>(MNCN)</td>
<td>H6</td>
<td>S1 (3)</td>
</tr>
<tr>
<td>Kernouvé (Kn)</td>
<td>(Smithsonian)</td>
<td>H6</td>
<td>S1 (3)</td>
</tr>
<tr>
<td>Portales Valley (PV)</td>
<td>USNM 6975 (SI)</td>
<td>H6</td>
<td>S1 (3)</td>
</tr>
<tr>
<td>Herdjaz (Hdj)</td>
<td>MNHN 2132 (MNHN-Paris)</td>
<td>L3.7</td>
<td>S4 (2)</td>
</tr>
<tr>
<td>Bjürbole (Bjb)</td>
<td>USNM 6705 (SI)</td>
<td>L4</td>
<td>S1 (5)</td>
</tr>
<tr>
<td>Tourine la Grosse (TIG)</td>
<td>MNHN 191 (MNHN-Paris)</td>
<td>L6</td>
<td>S3 (5)</td>
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<tr>
<td>Alfianello (Alf)</td>
<td>MNHN 885 (MNHN-Paris)</td>
<td>L6</td>
<td>S5 (2)</td>
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<td>Vouillé (Vll)</td>
<td>MNHN 148 (MNHN-Paris)</td>
<td>L6</td>
<td>S5 (4)</td>
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<td>Granes (Grn)</td>
<td>MNHN 2852 (MNHN-Paris)</td>
<td>L6</td>
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<td>Parnalle (Pr)</td>
<td>USNM 1613 (SI)</td>
<td>LL3.6</td>
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<tr>
<td>NWA 118 (N118)</td>
<td>Private collection</td>
<td>LL6</td>
<td>S3 (7)</td>
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</table>


Table 2. The Ge isotopic compositions of standards and geostandards, cleaning tests on San Carlos olivines for silicate purification, and BIR-1 digestion in Bola Teflon bombs at 150°C and 25 bar. Germanium concentrations were measured using an ICP-MS X-Series at CRPG, with a precision of less than 2%. \( n \) is the number of replicates.

<table>
<thead>
<tr>
<th></th>
<th>Ge ppm</th>
<th>( \delta^{72/70}\text{Ge} ) ± 2SD</th>
<th>( \delta^{73/70}\text{Ge} ) ± 2SD</th>
<th>( \delta^{74/70}\text{Ge} ) ± 2SD</th>
<th>( n )</th>
</tr>
</thead>
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<tr>
<td><strong>In-house standard</strong> Magura (IAB iron)</td>
<td>427(^{(1)})</td>
<td>0.42 0.06</td>
<td>0.58 0.11</td>
<td>0.78 0.08</td>
<td>9</td>
</tr>
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<td><strong>Standard solutions</strong> JMC-Ge</td>
<td>–0.15 0.08</td>
<td>–0.19 0.13</td>
<td>–0.32 0.11</td>
<td>9</td>
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<td></td>
<td>Aldrich-Ge</td>
<td>–1.01 0.07</td>
<td>–1.52 0.13</td>
<td>–1.97 0.12</td>
<td>9</td>
</tr>
<tr>
<td><strong>Geostandards</strong> BIR-1 Hotplate digestion</td>
<td>1.52(^{(2)})</td>
<td>0.32 0.04</td>
<td>0.42 0.07</td>
<td>0.59 0.06</td>
<td>7</td>
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<tr>
<td></td>
<td>Bomb digestion</td>
<td>1.50 0.31</td>
<td>0.42 0.14</td>
<td>0.59 0.01</td>
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<td>San Carlos Olivines Un-cleaned</td>
<td>0.89 0.35</td>
<td>0.48 0.08</td>
<td>0.70 0.08</td>
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<tr>
<td></td>
<td>Cleaned</td>
<td>0.90 0.36</td>
<td>0.47 0.14</td>
<td>0.71 0.09</td>
<td>3</td>
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</table>

\(^{(1)}\)Value from Luais (2007). \(^{(2)}\)Value from Luais (2012).
### H ordinary chondrites

#### Bulk

<table>
<thead>
<tr>
<th>Samples</th>
<th>Ni (wt%)</th>
<th>Co (ppm)</th>
<th>Co/N</th>
<th>Ge (ppm)</th>
<th>δ²⁸⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>δ²³⁵⁷⁷Ge %o</th>
<th>± 2SD</th>
<th>δ²⁴⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>N</th>
</tr>
</thead>
<tbody>
<tr>
<td>Dhajala (H3.8)</td>
<td>1.22</td>
<td>850</td>
<td>0.07</td>
<td>15.2</td>
<td>−0.16</td>
<td>0.06</td>
<td>−0.25</td>
<td>0.12</td>
<td>−0.37</td>
<td>0.15</td>
<td>8</td>
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<tr>
<td>Sharps (H3.4)</td>
<td>1.26</td>
<td>668</td>
<td>0.05</td>
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#### Metal

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<th>Co/N</th>
<th>Ge (ppm)</th>
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<th>± 2SD</th>
<th>δ²³⁵⁷⁷Ge %o</th>
<th>± 2SD</th>
<th>δ²⁴⁷⁰Ge %o</th>
<th>± 2SD</th>
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<td>0.04</td>
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<tr>
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<td>−0.57</td>
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<td>0.07</td>
<td>−0.35</td>
<td>0.1</td>
<td>−0.51</td>
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#### Silicate

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<th>Co (ppm)</th>
<th>Co/N</th>
<th>Ge (ppm)</th>
<th>δ²⁸⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>δ²³⁵⁷⁷Ge %o</th>
<th>± 2SD</th>
<th>δ²⁴⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>N</th>
</tr>
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<tbody>
<tr>
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<td>0.5</td>
<td>5</td>
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<td>−0.32</td>
<td>0.1</td>
<td>−0.53</td>
<td>0.07</td>
<td>−0.66</td>
<td>0.17</td>
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<td>−0.58</td>
<td>0.05</td>
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<td>−0.27</td>
<td>0.06</td>
<td>−0.42</td>
<td>0.11</td>
<td>−0.55</td>
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#### Sulfide

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<th>Ge (ppm)</th>
<th>δ²⁸⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>δ²³⁵⁷⁷Ge %o</th>
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<td>0.002</td>
<td></td>
<td></td>
<td></td>
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#### Chondrules: Dhajala

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<th>Co (ppm)</th>
<th>Co/N</th>
<th>Ge (ppm)</th>
<th>δ²⁸⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>δ²³⁵⁷⁷Ge %o</th>
<th>± 2SD</th>
<th>δ²⁴⁷⁰Ge %o</th>
<th>± 2SD</th>
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<td>−0.27</td>
<td>0.05</td>
<td>−0.28</td>
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<td>&gt;300µm fraction</td>
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<td></td>
<td></td>
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<td>0.06</td>
<td>−0.32</td>
<td>0.03</td>
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#### L ordinary chondrites

#### Metal

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<th>Co (ppm)</th>
<th>Co/N</th>
<th>Ge (ppm)</th>
<th>δ²⁸⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>δ²³⁵⁷⁷Ge %o</th>
<th>± 2SD</th>
<th>δ²⁴⁷⁰Ge %o</th>
<th>± 2SD</th>
<th>N</th>
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<td>Hedjaz (L3.7)</td>
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<td>104.3</td>
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<td>−0.3</td>
<td>0.11</td>
<td>−0.35</td>
<td>0.09</td>
<td>3</td>
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</table>

### Notes

1) Includes Ni isotopes.
2) Includes Ni isotopes.
Table 3. The Ni, Co, and Ge concentrations, as well as the $\delta^{72/70}$Ge, $\delta^{73/70}$Ge, and $\delta^{74/70}$Ge values of the ordinary chondrites relative to NIST 3120a. Reproducibility is given as 2SD. N is the number of replicates for $\delta^{72/70}$Ge, $\delta^{73/70}$Ge, and $\delta^{74/70}$Ge. Error of less than 5% for the Ge, Ni, and Co elemental compositions.

1. Dhajala was excluded from the mean due to its unusual Ge isotopic composition, see section 4.3 for explanations.  
2. There is no standard deviation for sulfide because only one measurement was done per sample due to their low concentration.
Table 4. The germanium isotopic factor among the bulk, metal, silicate, and sulfide of H ordinary chondrites, as well as the Ge elemental fractionation between metal and silicate.

<table>
<thead>
<tr>
<th>Samples</th>
<th>$\Delta^{74/70} \text{Ge}_{\text{bulk-metal}}$</th>
<th>$\Delta^{74/70} \text{Ge}_{\text{bulk-silicate}}$</th>
<th>$\Delta^{74/70} \text{Ge}_{\text{metal-silicate}}$</th>
<th>$\Delta^{74/70} \text{Ge}_{\text{bulk-sulfide}}$</th>
<th>$D\text{Ge}_{\text{metal-silicate}}$</th>
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<td>0.33</td>
<td>−</td>
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<td>0.19</td>
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<td>−</td>
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<td>$0.22 \pm 0.36$</td>
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Table 5. The oxygen isotopic compositions and percentage of fayalite (%Fa) in ordinary chondrites.

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<th>%Fa</th>
<th>Ref</th>
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<th>δ17O (‰)</th>
<th>Δ17O (‰)</th>
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| IIE Iron meteorites |     |     |           |           |           |      |     |
| Netschaévo        | 3.46|     | 2.37      | 0.57      |           | 10   |     |
| Miles             | 4.75|     | 3.04      | 0.57      |           | 10   |     |
| Weekeroo station  | 4.11|     | 2.58      | 0.62      |           | 10   |     |

References: 1: Clayton et al. (1991); 2: Folco et al. (2004); 3: This study; 4: Rubin (1990); 5: Dodd (1968); 6: Dodd & Jarosewich (1979); 7: Reisener et al. (2006); 8: Grossman & Zipfel (2001); 9: Mason (1963); 10: Clayton & Mayeda (1996)
Declaration of interests

☒ The authors declare that they have no known competing financial interests or personal relationships that could have appeared to influence the work reported in this paper.

☐ The authors declare the following financial interests/personal relationships which may be considered as potential competing interests: