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1 **TITLE**

2 Helium isotope evidence for cryptic magmatism along Central France

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13 **ABSTRACT**

14 Determining the source(s) of helium in regions of the crust where ³He anomalies occur
15 can be challenging when no surface manifestations of magmatic activity or a clear
16 active extension regime are present. This is the case of the Paris basin (France), where
17 geothermal fluids and oils ³He/⁴He have been previously shown to range from 0.02 R_a
18 up to 0.14 R_a (Marty et al., 1993; Pinti & Marty, 1995, 1998), whereas natural gases to
19 the south of the Paris Basin (near the Massif central) show unambiguous evidence for
20 mantle helium contribution, with ³He/⁴He up to 6.4 R_a (Bräuer et al., 2017). However,
21 higher ³He/⁴He than the crustal endmember (~ 0.01 R_a), as observed in the Paris Basin,
22 could be a sign of either (i) a weak contribution of mantle He (enriched in primordial
23 ³He) or (ii) the contribution of ³He naturally produced by nuclear reactions of lithium,
24 for which commercially exploitable concentrations have been reported in rare-metal-
25 rich granites. One way to distinguish between these two possibilities is to combine He
26 isotope systematics with other noble gas tracers of mantle inputs, such as xenon
27 isotopes. Here, we report the isotope compositions of helium and xenon, measured by
28 static and dynamic mass spectrometry, respectively, in gas samples collected in the
29 Nièvre County, between the Massif Central and the Paris Basin. ³He/⁴He ratios range
30 between 0.16 and 0.22 R_a. Combined with previous studies, our data are consistent with

31 a southward increase of He isotope ratios along major fault systems from the Paris
32 Basin ($\sim 0.02 R_a$) towards the Massif Central ($6.4 R_a$). This geographical trend, paired
33 with high precision Xe isotopes exhibiting an excess of mantle ^{129}Xe , shows a clear
34 mantle input that is most likely restricted to the major fault systems in Central France.
35 The occurrence of mantle-derived helium and xenon in Central France highlights the
36 presence of cryptic magmatism in areas with no apparent volcanic manifestations.

37 **Keywords:** noble gases; helium isotopes; xenon isotopes; mantle input; CO_2 reservoir

38 1 Introduction

39 On Earth, the two helium isotopes, ^3He and ^4He , are found in the atmosphere,
40 crust, and mantle in different proportions. ^3He is mainly of primordial origin, occurring
41 in higher proportions in the mantle, where it has been able to be preserved since Earth's
42 accretion (Craig & Lupton, 1976). ^4He on the other hand is primarily of radiogenic
43 origin and is concentrated in the continental crust (Ballentine & Burnard, 2002). The
44 production of ^4He primarily occurs by α -decay of $^{235,238}\text{U}$ and ^{232}Th , and is thus
45 dependent on the concentration of these two parent elements (Ballentine & Burnard,
46 2002). However, production of ^3He can also occur in the crust produced through the
47 thermal neutron capture on ^6Li by the reaction $^6\text{Li} (n, \alpha) ^3\text{H} (\beta^-) \rightarrow ^3\text{He}$ (Morrison &
48 Pine, 1955).

49 The $^3\text{He}/^4\text{He}$ ratio is thus an essential tool for determining the origin of different
50 gas reservoirs. In some regions of the crust, the interpretation of $^3\text{He}/^4\text{He}$ can be
51 ambiguous, such as in the Paris Basin (France), where no magmatic manifestations at
52 the surface, or evidence for modern tectonic extension are present, yet helium isotope
53 ratios show higher $^3\text{He}/^4\text{He}$ than expected for the continental crust (up to $0.14 R_a$; Marty
54 et al., 1993; Pinti & Marty, 1995, 1998). Helium isotopes have been investigated in
55 groundwater and crude oil from the Paris Basin and in gas manifestations in the Massif
56 Central, respectively (Bräuer et al., 2017; Marty et al., 1993; Pinti & Marty, 1995,
57 1998). While there is unequivocal evidence for a mantle contribution in the Massif
58 Central, with $^3\text{He}/^4\text{He}$ ratios up to $6.4 R_a$ (where R_a is the atmospheric $^3\text{He}/^4\text{He} = 1.39 \times$
59 10^{-6} ; Graham, 2002), a mantle contribution in the Paris Basin is less evident. Helium
60 isotope ratios ($0.02 - 0.11 R_a$, reaching $0.14 R_a$ for basement-hosted groundwater) are
61 close to the crustal end-member ($0.001 - 0.02 R_a$) as computed from average U+Th/Li
62 abundances in related lithologies of the Paris Basin (Pinti & Marty, 1998). Because
63 $^3\text{He}/^4\text{He}$ ratios of geothermal fluids and oils are higher than values predicted for a purely

64 crustal origin, Pinti and Marty (1998) proposed a mantle contribution to gases hosted by
65 basinal fluids. In agreement with this possibility, it was observed that the $^3\text{He}/^4\text{He}$ ratios
66 increase with depth in the basin, suggesting injection of mantle-derived helium from the
67 basement underlying the sedimentary units, and progressive dilution of this signal by
68 radiogenic, ^4He -rich helium produced in the sedimentary pile (Marty et al., 1993).
69 Helium isotope ratios also tend to be higher towards the southern margin of the basin
70 (0.09 Ra in groundwater and 0.14 Ra at the basement; Pinti and Marty, 1998). This
71 trend is possibly related to the North-South major fault system (“Sillon Houiller”) in
72 Central France, and to the addition of mantle-derived ^3He from the nearby Rhine rift
73 magmatism at the eastern border of the Paris Basin in Lorraine (Marty et al., 2003).
74 Alternatively, $^3\text{He}/^4\text{He}$ ratios higher than expected for the purely crustal end-member
75 could reflect local enrichments of lithium-derived ^3He , with He isotope variations
76 potentially tracking the occurrence of Li-rich lithologies of economical relevance.

77 Recently, a CO_2 -rich gas reservoir with He concentration reaching 0.04% has
78 been discovered while conducting helium exploration in the Nièvre County, between
79 the Massif Central to the south and the southern border of the Paris Basin to the north
80 (Fig. 1). Although not reaching commercial concentrations for helium ($> 0.3\%$), its
81 helium content is about 100 times higher than the atmosphere ($\sim 0.0005\%$; Glueckauf,
82 1946). This gas accumulation has the potential to shed light into the nature of the
83 intermediate $^3\text{He}/^4\text{He}$ ratios, between typical mantle values and the crustal range,
84 observed in the Paris Basin (e.g., Pinti and Marty, 1998). The geochemical description
85 of this reservoir is still at an early stage, and so our paper focuses on the use of isotope
86 geochemistry to identify the source(s) of the Nièvre gas.

87 In the present study, we sampled bulk gas samples using two distinct approaches:
88 copper tubes (dedicated to He isotope measurements) and Giggenbach bottles
89 (dedicated to Ar-Ne-Xe isotope measurements). The samples consist in free gases from
90 two CO_2 -rich wells (SP4 and SP7) and bubbling springs from the surrounding areas –
91 Gelin, Fonts-Bouillants, Fontaine des Vertus, Pougues-les-Eaux, Fontaines Salées, and
92 Bourbon-Lancy (Fig. 1). Most springs were cold ($12\text{-}14^\circ\text{C}$; Boineau & Maisonneuve,
93 1972; Risler, 1974), apart from Bourbon-Lancy, which has a temperature of 52°C
94 (Batard et al., 1982). To identify the different sources of helium, we analysed $^3\text{He}/^4\text{He}$
95 and $^4\text{He}/^{20}\text{Ne}$ on gas collected from all sites. We also analysed the major gas

96 composition, stable isotopes of CO₂ ($\delta^{13}\text{C}$) and N₂ ($\delta^{15}\text{N}$), and Ar, Ne, and Xe isotopes
97 for the SP4 and SP7 well gas samples.

98 **2 Geological setting**

99 The study area is located between the southern border of Paris Basin and the north
100 of the Massif Central and Limagne basin, west of the Morvan massif (Fig. 1). The
101 Massif Central and Morvan belong to the Hercynian belt. The basement of the basin is
102 made up of Hercynian granites, migmatites, gneisses and mica schists that outcrop in
103 the Decize horst (Clozier and Turland, 1982; Roger et al., 2010). The study zone is also
104 likely to contain uranium-rich carboniferous and permian deposits, mostly alternating
105 clay and sandstone, that formed during the dismantling of the Variscan belt under the
106 Mesozoic and Cenozoic cover (Beccaletto et al., 2015; Farjanel, 1989; Mercuzot et al.,
107 2021).

108 The area underwent magmatic and volcanic activity during the Hercynian period
109 and Paleogene (Baptiste, 2016; Farjanel, 1989). The emplacement of the Decize and
110 Montmarault granitic intrusion (321 ± 2 Ma U-Th-Pb dating on monazite; Joly, 2007)
111 and rare metal-enriched granites in the Morvan and Massif Central took place as the belt
112 began dismantling (Gloaguen et al., 2018; Roger et al., 2010). The most recent volcanic
113 activity occurred during the Paleogene, at ~90 km to the South of the Paris Basin, in
114 “Chaîne des Puys” related to the Alpine orogeny and the Cenozoic extension regime
115 (Barruol & Granet, 2002). The origin of volcanism in the area is still debated (Nehlig et
116 al., 2003). On the one hand, it could reflect the surface expression of a mantle plume
117 (Coisy & Nicolas, 1978; Froidevaux et al., 1974), as potentially suggested by seismic
118 tomography (Mazabraud et al., 2005), originating from the asthenosphere related to the
119 activity of a hot spot (Froidevaux et al., 1974) or to a passive rift (Merle & Michon,
120 2001). On the other hand, volcanism may result from crustal extension and melting of
121 the lithospheric material (Gautheron et al., 2005). One last possibility is that a deep
122 plume led to the formation of shallower plumes in the Sub Continental Lithospheric
123 Mantle (SCLM) and/or in the old oceanic lithosphere crust (~400 Ma subduction
124 during the Hercynian orogeny) (Buikin et al., 2005; Gautheron et al., 2005). Noble gas
125 data from the Massif Central, namely Xe and Ne isotopes, indicate a MORB-like
126 source, discarding the influence of a mantle plume (Moreira et al., 2018).

127 Seismicity is not restricted to the Massif Central and seismic events have been
128 recorded along the major structures of the area (Marche and Sillon Houiller faults).

129 Most faults appear to be related to the complex geodynamic history of the area
130 (Variscan subduction, collision, and exhumation) and are deeply embedded in the
131 basement, such as the Sillon Houiller and Loire faults (Baptiste, 2016; Regorda et al.,
132 2020; Roger et al., 2010; Vanderhaeghe et al., 2020). The Magnetic Paris Anomaly
133 follows the same N-S orientation as the Loire fault in the North of Massif Central. Even
134 if its origin is debated, this anomaly could be linked to a deep major structure, such as a
135 main strike-slip fault or an oceanic suture covered by the Paris Basin sediments
136 (Baptiste, 2016). Variscan faults oriented N20-30 have been reactivated into shear and
137 then normal faults during the formation of the Carbonifereous pull-out and Permian
138 hemi-graben (Beccaletto et al., 2015; Farjanel, 1989). They were also reactivated by
139 extension and compression at the end of the Jurassic (E-W) during the opening of the
140 Tethys, amid the Cretaceous (SW-NE extension and N-S compression) and the W-E
141 extension in the Tertiary (Guillocheau et al., 2000; Jolivet et al., 2021; Manatschal et
142 al., 2021; Roger et al., 2010). Gas-rich springs and drillings are located along faulted
143 areas with mainly N-S directions. SP4, SP7, Gelin, Fonts-Bouillants, and Fontaine des
144 Vertus follow a single fault, the Saint-Parize-le-Châtel fault while Pougues-les-Eaux is
145 located in an area affected by multiple normal faults. Bourbon-Lancy and Fontaines
146 Salées are situated along Morvan bounding fault.

147 **3 Methods**

148 *3.1 Bulk gas compositions*

149 Gases occur either as bubbles in springs (Gelin, Fonts-Bouillants, Fontaine des
150 Vertus, Pougues-les-Eaux, Bourbon-Lancy and Fontaines Salées) or as free CO₂-rich
151 gas in two drill holes (SP4 and SP7). Samples were collected during three field
152 campaigns and/or well testing operations, in August-September, 2021, February and
153 May-June, 2022 (Table 1). SP4 and SP7 drill holes are rooted in the Triassic sediments
154 and cross the Saint-Parize-le-Châtel fault. Gases were sampled in evacuated vials or
155 Tedlar Bags after purging the tubes connecting them to the wellheads. The springs near
156 SP4 (Gelin, Fonts-Bouillants, Pougues-les-Eaux, and Fontaine des Vertus) were
157 sampled before and after the well test to detect potential variations in helium isotopes.
158 Bubbling gases from springs were collected through funnel and Teflon tubing in
159 previously evacuated vials (Pasquet et al., 2021) or Tedlar bags (Battani et al., 2010).
160 Finally, the gas composition was measured by gas chromatography (GC)
161 (Supplementary material; Pasquet et al., 2021).

162 3.2 *Noble gases*

163 3.2.1 $^3\text{He}/^4\text{He}$ and $^4\text{He}/^{20}\text{Ne}$

164 The analyses of $^4\text{He}/^{20}\text{Ne}$ and $^3\text{He}/^4\text{He}$ ratios of gases sampled in copper tubes
165 were performed, respectively, on a MKS Microvision 2 quadrupole mass spectrometer
166 (QMS) and a split flight tube noble gas mass spectrometer (Helix SFT - Thermo Fisher
167 Scientific) at the Centre de Recherches Pétrographiques et Géochimiques (CRPG) noble
168 gas analytical facility (Mabry et al., 2013). The standard deviation of the $^3\text{He}/^4\text{He}$ ratios
169 ranged from 1.4 to 3.7 % based on the replicate analysis of standard He (atmospheric).
170 Results are corrected for blanks, which contributed < 0.1 % of the total He abundances
171 (Table 2).

172 3.2.2 *Ar, Ne, and Xe isotopes*

173 Gases were collected in 1.5 L Giggenbach bottles (Giggenbach & Goguel, 1988;
174 Bekaert et al. 2019) containing 4N NaOH solution in which CO_2 and other reactive
175 species (e.g., S compounds) can be trapped. This method permits the concentration of
176 inert gases (mainly N_2 and noble gases and possible traces of CH_4 , CO , H_2 , etc) in the
177 vacuum headspace above the solution. In the laboratory, the gaseous fraction was
178 equilibrated under vacuum with an evacuated metal bottle fitted with two metal valves
179 (Nupro SS4H[®]), where the gas could be stored for long periods of time without
180 diffusive leakage or exchange with atmospheric gases. Aliquots of gas were expanded
181 in a vacuum line and the abundances of He and Ar as well as Ar isotopic ratios were
182 measured after purification via static mass spectrometry on a GV Helix MC mass
183 spectrometer. Two aliquots of the SP4 well gas were also sampled from the metal
184 bottles and analysed for all noble gases (Table 3). Gases were purified using hot
185 (600°C) and cold (25°) Ti-getters and, after cryogenic separation (Broadley et al.,
186 2022), noble gas abundances and isotopic ratios were analysed with a Helix MC+ mass
187 spectrometer at the Centre de Recherches Pétrographiques et Géochimiques (CRPG)
188 noble gas analytical facility (Bekaert et al., 2019; Broadley et al., 2020).

189 At last, argon and xenon isotope ratios were measured in the Seltzer Lab at
190 Woods Hole Oceanographic Institution via dynamic isotope-ratio mass spectrometry
191 (Bekaert et al., 2023; Seltzer & Bekaert, 2022) for one sample of the SP7 well gas
192 collected in a large Giggenbach bottle (1.5 L) following the sampling procedure detailed
193 by Bekaert et al. (2019). With a dynamic mass spectrometer, often called an “isotope
194 ratio mass spectrometer,” the rapid switching between analysis of sample and reference

195 gas streams reduces the effect of instrumental drift (McKinney et al., 1950) and allows
196 for many comparable measurements of purified sample gas and standard gas of known
197 isotopic composition over the course of an analysis, hence ensuring accurate and
198 unbiased isotopic ratio determination at the sub-per mil level. Results are in permil (‰)
199 and notated as δ following $\delta R = [(R_{sample}/R_{atmosphere}) - 1] \times 1000$; where R is a
200 Xe isotope ratio. Raw measurements were corrected for diffusive transport fractionation
201 (DTF; light isotope enrichment due to physical subsurface fractionation; Bekaert et al.,
202 2023) using $\delta^{128}\text{Xe}/^{130}\text{Xe}$, yielding a mass-dependent correction of $0.504 \pm 0.157 \text{‰ u}^{-1}$.
203 Our reported uncertainties on Xe isotope data reflect both analytical and DTF-correction
204 sources of uncertainty (Table 3).

205 3.3 *Stable isotopes*

206 The carbon and nitrogen isotopic compositions of CO₂ and N₂, respectively, were
207 measured with a Thermo Scientific MAT 253 Mass Spectrometer at the Centre de
208 Recherches Pétrographiques et Géochimiques (CRPG) stable isotopes analytical
209 facility. Sampling was performed using Giggenbach bottles and evacuated glass bottles.

210 3.3.1 $\delta^{13}\text{C}$

211 For the $\delta^{13}\text{C}$ measurements of the CO₂, CO₂ and H₂O were trapped into liquid
212 nitrogen and then cryogenically separated. All values are reported in the delta notation
213 in permil (‰) relative to PDB, with error below 0.2 ‰ (Table 2).

214 3.3.2 $\delta^{15}\text{N}$

215 Nitrogen isotope measurements were done using the non-condensable fraction of
216 the Giggenbach bottles, which consists mainly of N₂, CO, and noble gases since CO₂
217 has been removed. Both CO and N₂ gases were transferred in a sealed pyrex tube
218 containing CuO and Cu₂O grains to be oxidized into N₂ and CO₂ at 450 °C for 1 hour.
219 Prior to being used, the CuO and Cu₂O grains have been heated few minutes under
220 vacuum at 350°C and pumped overnight at room temperature to remove adsorbed gases.
221 The produced CO₂ was then cryogenically trapped and N₂ was collecting using a
222 Toepler pump. All values are reported using the delta notation, in permil (‰) relative to
223 air, with error ~0.5 ‰ calculated from the reproducibility of duplicates (Table 2).

224 4 Results

225 4.1 Wells

226 During the main flow of the SP4 well test, which lasted approximately 10h,
227 samples were collected for $^3\text{He}/^4\text{He}$ and $^4\text{He}/^{20}\text{Ne}$ measurements every hour. Sample
228 names follow the format “site name-sampling order + copper tube number” and
229 duplicate samples have an “M1” at the end; sampling order is “100” when before well
230 test, “200” during, and “300” if sampled after the well test (Table 1). For the SP7 well,
231 only one sample was collected for noble gas measurements. SP4 and SP7 are CO_2 -
232 dominated, consisting of $\sim 94\%$ CO_2 and 6% N_2 and only traces of CH_4 ($< 0.6\%$);
233 helium contents range from 394 to 426 ppm vol. The $\delta^{13}\text{C}$ value of the SP4 sample is -
234 10.1 ‰, which is compatible with a primarily organic source of CO_2 or non-marine
235 carbonates (Table 2).

236 Throughout the SP4 well test, no significant change on helium abundance or
237 isotopic ratios was detected. $^3\text{He}/^4\text{He}$ ratios range from 0.17 to 0.21 (± 0.01) R_a , with no
238 temporal variation (Table 2). $^4\text{He}/^{20}\text{Ne}$ ratios are consistently > 2000 , however,
239 measurements are not precise due to the low content of ^{20}Ne relative to ^4He , with the
240 former showing values very close to the baseline. During the SP7 well test, we only
241 collected one sample, which shows similar results with a $^3\text{He}/^4\text{He}$ ratio of 0.18 (± 0.01)
242 R_a and $^4\text{He}/^{20}\text{Ne} > 2000$.

243 The $^{40}\text{Ar}/^{36}\text{Ar}$ ratios, measured for two aliquots of the SP4 well gas, are 2866 ± 5
244 and 2574 ± 11 , respectively, with the $^{38}\text{Ar}/^{36}\text{Ar}$ ratios being indistinguishable from
245 atmosphere (Table 3). The $^{20}\text{Ne}/^{22}\text{Ne}$ (9.40 ± 0.05) and the $^{21}\text{Ne}/^{22}\text{Ne}$ ratios ($0.0621 \pm$
246 0.0011) deviate slightly from the atmospheric Ne value, indicating contribution of
247 crustal (and possibly mantle) neon (see section 5.2). Analyses of SP7 show $^{40}\text{Ar}/^{36}\text{Ar}$
248 (696.6 ± 0.2) above the air value, but still lower than the SP4 well, indicating possible
249 atmospheric contamination during drilling operations. Xenon isotopic ratios of SP7 by
250 dynamic mass spectrometry (Table 3) indicate contribution of fissiogenic and mantle Xe
251 (see section 5.2).

252 4.2 Springs

253 Fontaines Salées and Bourbon-Lancy are N_2 -dominated (93 % and 97 %, respectively)
254 while Fontaine des Vertus is comprised of 63 % CO_2 and 39 % N_2
255 (Table 1). The remaining springs are CO_2 -dominated ranging from 95 to 97 %. For all
256 springs, only traces of CH_4 are identified ($< 0.6\%$). Helium contents range from 245 to

257 426 ppm vol for the CO₂-dominated sites, and from 2874 to 45794 ppm vol for Fontaine
258 des Vertus, Fontaines Salées, and Bourbon-Lancy.

259 The springs near SP4 (Gelin, Fonts-Bouillants, and Fontaine des Vertus) present
260 ³He/⁴He ratios similar to those of SP4, ranging from 0.16 to 0.21 (± 0.01) R_a, as well as
261 the Pougues-les-Eaux spring which is 25 km north from SP4 (0.22 ± 0.01 R_a) (Table 2).
262 Fontaines Salées and Bourbon-Lancy which are located 80 km NE and 50 km SE of
263 SP4, have distinct ³He/⁴He signatures (0.04 ± 0.00 and 0.03 ± 0.00 R_a, respectively) that
264 are close to the purely radiogenic end-member. ³He/⁴He variations do not correlate with
265 gas compositions, with similar ³He/⁴He for springs near SP4 (including Fontaine des
266 Vertus) despite much higher helium and N₂ contents.

267 δ¹³C of CO₂ values show little variation among the springs, ranging from -11.8 to
268 -10 ‰ (vs. PDB), except for Fontaines Salées (-20 ‰), which suggests either different
269 sources of CO₂ or different processing in the crust. Weinlich (2005) reported
270 comparable values of -13.9 ‰ for a N₂-dominated spring at Pougues-les-Eaux (74 %
271 N₂) and of -20.4 ‰ at Saint Père (94 % N₂), near Fontaines Salées. The Pougues-les-
272 Eaux spring in the present study is dominated by CO₂ (95 vol. %) and has a very similar
273 δ¹³C (-10.3 ‰) to the one described in Weinlich (2005). Moreover, the springs with
274 lower CO₂ contents tend to present lighter δ¹³C values (Fig. S1) that could indicate
275 chemical processes during transport in the crust due to solution of CO₂ in the water
276 phase and/or carbonate precipitation, concentrating helium and N₂ in the gas phase.
277 δ¹⁵N results for the springs from both groups display similar values (+6.1 to +8.4 ‰)
278 compatible with a primarily crustal source.

279 5 Discussion

280 5.1 Source of helium

281 In European basins of various ages, ³He/⁴He values range between 0.01 and 3.9
282 R_a, reaching ~0.1 R_a in the western Alps (Marty et al., 1992; Oxburgh et al., 1986).
283 Oxburgh et al. (1986) suggested that this wide range in Neogene basins was related to
284 their tectonic setting. Helium isotope compositions in the Pannonian basin reach values
285 up to 3.9 R_a consistent with an extension regimen and clear mantle input while the
286 Molasse and Po basins show a range of ³He/⁴He between 0.02 and 0.09 R_a (Marty et al.,
287 1992; Oxburgh et al., 1986), which, for the Molasse basin, Oxburgh et al. (1986)
288 interpret as being of crustal origin despite the apparent small mantle contribution (up to
289 0.9%). In the Paris Basin, ³He/⁴He ratios are close to the range of values for the Molasse

290 and Po basins (0.02 to 0.11 R_a), reaching 0.14 R_a at the basement in Couy (Pinti &
291 Marty, 1998). These values are higher than the theoretical crustal $^3\text{He}/^4\text{He}$ ratio (0.001-
292 0.02 R_a), based on the U-Th and Li values of the underlying sedimentary rocks. These
293 elevated $^3\text{He}/^4\text{He}$ values were therefore considered to be evidence for the input of
294 mantle material with a higher than crustal $^3\text{He}/^4\text{He}$ (Pinti & Marty, 1998). However,
295 with only small $^3\text{He}/^4\text{He}$ variations, up to 0.1 R_a , it is difficult to fully discriminate this
296 component from variation in the crustal $^3\text{He}/^4\text{He}$ ratio. In addition, these areas show no
297 evidence of igneous manifestations, and there is the possibility that ^3He could be
298 produced in the crust without mantle input (see section 5.3.1).

299 In the $^3\text{He}/^4\text{He}$ versus $^4\text{He}/^{20}\text{Ne}$ plot (Fig. 2), we identify two groups of samples:
300 one of them contains Fontaines Salées and Bourbon-Lancy, and the other one contains
301 SP4, SP7, surrounding springs, and Pougues-les-Eaux. Considering the typical values
302 found in the various basins in Europe, the $^3\text{He}/^4\text{He}$ signature found at Fontaines Salées
303 and Bourbon-Lancy is within the European crustal range (0.02 – 0.1 R_a). The remaining
304 samples show significantly higher $^3\text{He}/^4\text{He}$ (0.16 – 0.22 R_a). However, in any case, the
305 main source of helium in the basin is from the continental crust.

306 Using U, Th, and Li contents of lithologies from the Paris Basin (Pinti & Marty,
307 1998) (since rock chemistry data is not available for our target area), we calculated the
308 theoretical $^3\text{He}/^4\text{He}$ signature to be between 0.001 and 0.016 R_a , based on radiogenic
309 (^4He) and nucleogenic (^3He) production equations (Eq. 1; Andrews, 1985). For the SP4
310 sample group that includes SP4, SP7, nearby springs, and Pougues-les-Eaux, we
311 compute that the higher $^3\text{He}/^4\text{He}$ ratios can either be accounted for by, (i) either
312 nucleogenic ^3He production from the presence of high amounts of lithium in the host
313 rocks, (ii) or by a 2.5 % mantle input to the original crustal signature. In the next
314 sections, we discuss both possibilities for the $^3\text{He}/^4\text{He}$ signature we find for the SP4
315 group.

316 (Eq. 1) $P_3/P_4 = f(R, Na, Mg, Al, Si, Ca)F_{Li}/(3.2108 \times 10^6 + 7.7633 \times 10^5 R)$

317 Where P_3/P_4 is the theoretical $^3\text{He}/^4\text{He}$ based on the ^3He and ^4He production
318 rates, $f(\dots)$ is the neutron production rate based on the composition, F_{Li} is the fraction of
319 neutrons that are captured by ^6Li in the rock matrix, and R is the Th/U ratio.

320 5.1.1 Nucleogenic ^3He production

321 As previously stated, one potential production pathway to increase the $^3\text{He}/^4\text{He}$
322 from crustal values to that measured in the SP4, SP7, surrounding springs, and Pougues-
323 les-Eaux samples is through the localised production of nucleogenic helium (^3He) via
324 thermal neutron capture on ^6Li via the reaction $^6\text{Li}(n,\alpha)^3\text{H}(\beta^-) \rightarrow ^3\text{He}$ (Morrison & Pine,
325 1955). To raise the $^3\text{He}/^4\text{He}$ signatures of the SP4 sample group (Fig. 2, Table 2) from
326 an assumed original crustal signature (0.001 – 0.016 R_a) we calculate that it would
327 require Li concentrations in the basement and/or host rocks to be on the order of 600
328 ppm (Andrews, 1985; Eq. 1). This calculation assumes that all the excess ^3He within
329 the samples is produced from the thermal neutron capture by ^6Li and that the basement /
330 host rock contains typical crustal U-Th concentrations (1.8 and 7.2 ppm, respectively;
331 Krauskopf and Bird, 1967). Although this amount of Li is atypical in view of the
332 average continental crust value (30 ppm Li), small proximal basins such as the ones in
333 the Massif Central have the potential to accumulate lithium (Sun et al., 2022).

334 One of the major concentrated sources of lithium in the crust is surface and
335 subsurface brines. Near the study region (~ 100 km), in the Massif Central, a
336 geothermal brine with 81 ppm Li has been identified in the Limagne basin (Fonroche
337 geothermie, 2019) but values higher than that have not yet been reported. Brines from
338 the area were previously described with much lower lithium contents (e.g., Millot et al.,
339 2007; 7 ppm). In the Paris Basin, lithium concentration is low, in brines associated with
340 Triassic salt deposits (up to 60 ppm; Fontes & Matray, 1993a; Millot et al., 2011) and
341 oil fields (4.8 ppm; Fontes & Matray, 1993b b).

342 Apart from brines and geothermal water, the main *hard-rocks* that are sources of
343 lithium are rare-metal-rich granites such as the Beauvoir, Montebbras, and Chavence
344 granites, near our study area, LCT pegmatites, and greisen deposits produced by high-
345 temperature hydrothermal alteration of the aforementioned types (Gourcerol et al.,
346 2019). The lithium mineralizations in the Massif Central are usually located near
347 important ductile faults (Gloaguen et al., 2018).

348 Considering the various sources of lithium with production potential (Gloaguen et
349 al., 2018) in the Massif Central, near the study area, it is possible to consider that a Li-
350 rich granitic basement produces significant quantities of helium. The closest lithium
351 sources are the Chavence (56 km SE) and Beauvoir granites (72 km SW), the latter

352 containing lithium concentrations between 367 and 5,220 ppm in its upper units
353 (Raimbault et al., 1995). Taking the U, Th, and Li contents from the Beauvoir granite
354 (16.2 - 52.4, 1.15 - 5.63, and 367 - 5,220 ppm, respectively; Raimbault et al., 1995), the
355 theoretical $^3\text{He}/^4\text{He}$ ratios would range between 0.12 and 1.66 R_a . In that case, helium
356 could be produced in the basement, before migrating upwards to the basin and being
357 stored in the sedimentary layers, acting as a reservoir. Another possibility is that, as a
358 proximal basin, detrital material from the Li-rich rocks accumulates in the basin and ^3He
359 is produced in the sedimentary layers. The Chavence granite is located near Bourbon-
360 Lancy (~ 9 km) where we find a crustal signature for helium isotopes; however, at
361 Bourbon-Lancy, the spring is directly on top of the basement and not in a sedimentary
362 basin. To observe the same signatures as for the SP4 sample group at Bourbon-Lancy, a
363 sedimentary basin that accumulates and traps the produced gas is required, as otherwise
364 degassing from a Li-rich source would only occur locally and relatively far from the
365 spring. There are also non-conventional sources of lithium that have recently been
366 brought to light and that cannot be discarded, such as coal (e.g., Qin et al., 2015).
367 Considering the regional geology, Stephanian-Permian basins in the Massif Central
368 region have coal levels (Mercuzot et al., 2021) that could potentially provide the
369 required amounts of lithium.

370 In the context of local production, Martel et al. (1990) found $^3\text{He}/^4\text{He}$ ratios (0.025
371 R_a) higher than the expected theoretical value (0.0033 R_a) for the Carnmenellis granite-
372 hosted water (England). They explain this observation by presenting a model of
373 preferential release of ^3He from the rocks, relative to ^4He . However, this model only
374 works if some conditions are satisfied to retain ^4He within the minerals – i.e., U and Th
375 get concentrated in accessory phases and the host rock has remained at low temperature
376 since formation (hence limiting He loss through diffusion). In our current study, testing
377 this model would require detailed mineralogical characterization, but currently the
378 composition of underlying basement and host rock is not well known. Furthermore, the
379 Paris Basin has a relatively high heat flux (60 to 120 mW/m²; Dentzer et al., 2016;
380 Torgersen, 1993) which makes this model unlikely.

381 5.1.2 Mantle contribution

382 The alternative explanation for $^3\text{He}/^4\text{He}$ ratios above typical crustal values is a
383 contribution of mantle-derived ^3He . This could originate from an old magmatic source

384 in the crust, put in place around the time of basin formation, or from more recent
385 magma intrusions.

386 Méjean et al. (2020) proposed a model in which a fossil magmatic signal is
387 diluted over time by radiogenic ^4He production in the crust. This model is appropriate to
388 explain the mantle input observed in the Southern Quebec groundwater since a clear
389 alignment of Cretaceous magmatic intrusions is present (Monteregian Hills Igneous
390 Province). For the Paris Basin, however, no evidence of magma emplacement in such
391 scale during its formation is found.

392 Pinti and Marty (1998) describe two possibilities for the minor mantle signature in
393 the Paris Basin observed in He isotopes – either from Permo-Carboniferous intrusions
394 in Couy (lamprophyre dykes) or from a deep-seated magmatic source. The former could
395 explain localized variations in the $^3\text{He}/^4\text{He}$ ratios but is unlikely to explain the
396 signatures observed across a broad area. Furthermore, if the mantle input was coming
397 from magmatic bodies that have already crystallized, helium would need to be liberated
398 from minerals. For such a process to happen, certain temperatures must be reached to
399 release He from the mineral; in the case of olivine crystals, helium release from fluid
400 inclusions seems to start at $\sim 500^\circ\text{C}$ (Tolstikhin et al., 2010) while closure temperatures
401 are between 143 and 244°C (Shuster et al., 2004; Wang et al., 2015). In this case,
402 diffusion is the primary process, and will not only depend on the current isotherms but
403 also on the age of the crystallized body. Furthermore, cooling rates may play a role in
404 controlling the availability of He since, if the heat flux has been continuously high in
405 the region and the minerals crystallized long ago, the majority of the helium would
406 likely have been lost from the magmatic body. A deep-seated magmatic source could
407 explain the vertical $^3\text{He}/^4\text{He}$ gradient observed (Marty et al., 1993; Pinti & Marty, 1998)
408 from the basement-hosted groundwater with 0.12 - $0.14 R_a$, going through the Triassic
409 aquifer with $0.08 R_a$, and terminating at the Middle Jurassic waters at $0.02 R_a$.

410 In Fig. 3, we observe a geographical trend with $^3\text{He}/^4\text{He}$ decreasing from the
411 South (Massif Central) towards the north (Paris Basin) while the SP4 sample group
412 shows intermediate $^3\text{He}/^4\text{He}$ ratios. Considering that this trend coincides with the N-S
413 orientation of the fault systems, it could point to a localized weakening of the crust that
414 facilitates mantle input by degassing of a melt generated by decompression and partial
415 melting. The horizontal transport of gases with a mantle signature from a relatively
416 young magmatic system through faults is unlikely since the most recent active volcanic

417 system in the region is located in the Chaîne des Puys (4040 BCE), 90 km south of our
418 study area in the Massif Central (2.9- 3.2 Ra; Bräuer et al., 2017). Also, no evidence for
419 transport of gas via groundwater is observed in the dataset: transport through large
420 distances would entail gas-water interaction and the $^4\text{He}/^{20}\text{Ne}$ ratios are consistently
421 above 2000 and far from the air-saturated water value (~ 0.28). For the eastern segment
422 of the Paris Basin, further from our study area in the Lorraine region, Marty et al.
423 (2003) proposed that the elevated signatures (0.4 R_a) might originate from nearby rifting
424 in the Rhine graben (1.7 R_a ; Griesshaber et al., 1992), which is up to 120 km away from
425 their studied sites. Similarly, for the Drôme County, in the subalpine massifs (SE
426 France), $^3\text{He}/^4\text{He}$ ratios have a wide range and can reach relatively high values (0.02 to
427 3.64 R_a ; Marty et al., 1992) while being ~ 70 km away from the nearest volcanic
428 systems (Ardèche and Auvergne). In both studies, $^4\text{He}/^{20}\text{Ne}$ and $^3\text{He}/^4\text{He}$ ratios of
429 groundwater samples plot along a mixing line between ASW and a crust/mantle
430 component and thus water could work as the transport medium of the magmatic
431 $^3\text{He}/^4\text{He}$ component.

432 The input of mantle ^3He in some of our sample sites may therefore come directly
433 from the mantle underlying the continental crust. In the Pannonian basin and Rhine
434 graben, the crustal thickness is low, around 25 km and heat flow is on the order of 80
435 mW/m^2 (Čermák, 1982). Similarly, the current crustal thickness of the Paris Basin
436 varies between 29 and 41 km (Lefort & Agarwal, 1996; Prijac et al., 2000) while the
437 heat flux is between 60 and 120 mW/m^2 (Dentzer et al., 2016; Torgersen, 1993).
438 However, in the Morvan, the crustal thickness is lower at between 27 and 29 km (Lefort
439 & Agarwal, 1996) despite the typical crustal $^3\text{He}/^4\text{He}$ signatures (up to 0.04 R_a ; current
440 study). This suggests that crustal thickness and diffusive transport of mantle material
441 through the crust is unlikely to be the dominant source of mantle ^3He within these
442 systems. Therefore, Marty et al. (1993) interpret the anomalous $^3\text{He}/^4\text{He}$ signatures as
443 the result of degassing of partial melts at depth, which explains the vertical ^3He flux.

444 Although the commonly known processes by which mantle helium enters the
445 crust are via magmatic intrusions, degassing of magma, or diffusion, in areas without
446 recent volcanism and under either extension or compression regimes, mantle input can
447 also occur via advective flow controlled by faults due to improved permeability
448 (Kennedy & van Soest, 2007; Kulongoski et al., 2005, 2013). In the model proposed by
449 Kennedy & van Soest (2007), a vertical fault splay must penetrate the ductile lower

450 crust and reach the mantle acting as a channel for fluid flow. To satisfy this model,
451 strain localization must occur to generate vertical faults in depth; Kennedy & van Soest
452 (2007) propose this process is possible with increased dextral shear strain. The areas
453 where mantle input via advective flow have been described (Kennedy & van Soest,
454 2007; Kulongoski et al., 2005, 2013) are seismically active regions, however, which is
455 not the case for the Paris Basin. Slow deforming areas can present seismicity such as in
456 the Armorican Massif or the Massif Central, west and south of the Paris Basin,
457 respectively; the main difference between them, in a tectonic context, is the presence of
458 a thick unfaulted Meso-Cenozoic sedimentary pile that acts as a “patch” and protects the
459 basement rocks in the Paris Basin from brittle movement by improving the cohesion and
460 decreasing the thickness of the breakable portion of the upper crust (Petit et al., 2019).
461 Our study area, nonetheless, is at the border of the basin and contains a series of faults.
462 Further South, closer to the Massif Central, earthquakes with magnitudes up to M4.4
463 have been recorded (Mazabraud et al., 2005). The closest to the study area (~ 40 km
464 south) was an M4.1 in 1977, near Vieure (46.52°N, 2.93°E; Mazabraud et al., 2005). To
465 test if mantle input occurs via advective flow controlled by faults, a detailed structural
466 and geophysical study of the region is necessary.

467 In any case, the anomalous $^3\text{He}/^4\text{He}$ signatures – above typical crustal values for
468 this part of the crust – seem to be focused along the N-S fault systems with an apparent
469 decreasing gradient starting at the Massif Central volcanic area up to the Paris Basin.
470 This trend, coupled with the vertical $^3\text{He}/^4\text{He}$ gradient observed at the Paris Basin
471 $^3\text{He}/^4\text{He}$ (Marty et al., 1993; Pinti & Marty, 1998), seem to point to a localized and
472 limited mantle input along these faults, which is endorsed by the ^{129}Xe excess observed
473 at SP7 (see section 5.2).

474 5.2 *Ar, Ne, and Xe isotope constraints*

475 The Ne isotopic signature of SP4 is dominated by a nucleogenic production within
476 the continental crust, however, the data point lies slightly above the typical crust-air
477 mixing line (Fig 4), which could be due to either (i) simple two component mixing
478 between air and a Ne endmember between modern crustal values and Archean terrains,
479 or (ii) a limited contribution of mantle Ne to crustal and atmospheric Ne. The fact that
480 there is only a single data point does not allow us to draw mixing line(s) and therefore
481 precludes a definitive conclusion about a possible contribution of mantle Ne.

482 The argon isotopic composition of SP4 gas (Table 3) indicates either atmospheric
483 or mantle origin for the $^{38}\text{Ar}/^{36}\text{Ar}$ ratio and a significant contribution of radiogenic ^{40}Ar
484 produced by the decay of ^{40}K over long periods of time ($T_{1/2} = 1.25$ Ga). Indeed, the
485 $^{40}\text{Ar}/^{36}\text{Ar}$ ratios of 2866 are one order of magnitude higher than the atmospheric value
486 of 298.6. High $^{40}\text{Ar}/^{36}\text{Ar}$ values are observed in both crustal and mantle rocks and gases,
487 so it is not possible a priori to decipher the origin of such radiogenic Ar, but the
488 continental environment of the gas reservoir and the low $^3\text{He}/^4\text{He}$ ratio compared to
489 typical mantle He values suggests a strong crustal contribution.

490 At the level of ultra-high precision provided by dynamic mass spectrometry, the
491 isotopic composition of SP7 xenon is shown to deviate from that of atmospheric Xe.
492 High precision Xe analysis from SP7 (Fig 5) shows a small yet distinguishable ^{129}Xe
493 excess (relative to the atmosphere). ^{129}Xe excess is linked to ^{129}I decay either by ^{238}U
494 spontaneous fission, cosmogenic origin, or trapped in the mantle reservoir from
495 production in the first ~ 100 Myr of the Earth formation (Moreira, 2013;
496 Mukhopadhyay & Parai, 2019). Although ^{129}Xe can be produced in the crust by ^{129}I
497 decay from ^{238}U spontaneous fission, considering the $^{136}\text{Xe}/^{130}\text{Xe}$ is very low (corrected
498 absolute ratio of 2.207) and the upper limit for $^{129}\text{Xe}/^{136}\text{Xe}$ of ^{238}U -fissiogenic Xe is
499 0.001 (Eikenberg et al., 1993), the $^{129}\text{Xe}/^{130}\text{Xe}$ originated from fissiogenic ^{238}U would
500 be minimal ($< 1\%$).

501 Cosmogenic ^{129}I is concentrated in sediments enriched in carbon (Santos et al.,
502 2007); Holland et al. (2013) attributed the ^{129}Xe observed in the Timmins mine fluids to
503 cosmogenic ^{129}I due to the age of the host rocks/fluids (Archean), presence of carbon-
504 rich sediments, and absence of any evidence for mantle input since the $^3\text{He}/^4\text{He}$ ratios
505 were within typical crustal ranges (up to $0.01 R_a$). As discussed in section 5.1.1,
506 Stephanian-Permian basins in the Massif Central region have coal levels (Mercuzot et
507 al., 2021), however, the geology of the basin of our current study is not yet well
508 established. Thus, in our study area, considering the helium isotope signature above
509 typical crustal values (up to $0.22 R_a$), ^{129}Xe excess is linked to mantle input.

510 Notably, heavy Xe isotopes are clearly enriched relative to the lighter ones (e.g.,
511 ^{130}Xe). Such enrichment is likely to result from the contribution of fissiogenic Xe due to
512 the spontaneous fission of long-lived ^{238}U (Mukhopadhyay & Parai, 2019). Thus, the
513 isotopic composition of xenon confirms that SP7 hosts gas components derived from
514 both the mantle and crust.

515 5.3 *CO₂/³He and stable isotope constraints*

516 The overall $\delta^{13}\text{C}$ of CO₂ signatures are outside the MORB range ($-5 \pm 2 \text{ ‰}$) and
517 within range of non-marine carbonates, graphite, and marine plants (Favara et al.,
518 2002). However, for the SP4 group, with exception of Fontaine des Vertus, CO₂/³He is
519 close to the MORB values (~ 1 to 6×10^9 ; Marty & Zimmermann, 1999), ranging
520 from 8×10^9 to 1×10^{10} . In a CO₂/³He versus $\delta^{13}\text{C}$ plot (Sano and Marty, 1995; Fig. 6),
521 the CO₂-dominated samples show a mixing between three components, namely mid-
522 ocean ridge basalts (MORB), organic sediments, and marine limestones (or slab
523 carbonates). Considering that ³He is mainly of mantle origin, the end-member
524 contributions would range from 15 % to 30 % MORB, 40 % to 55 % limestones, and
525 25% to 30 % sediments contribution to the source of CO₂. Secondary processes can still
526 play an important role in the CO₂ and $\delta^{13}\text{C}$ values (Supplementary material); Gilfillan et
527 al (2009) propose that CO₂/³He variations in different gas fields are due to CO₂ loss
528 either from dissolution in groundwater or carbonate precipitation, and that the $\delta^{13}\text{C}$ will
529 vary according to the water pH in a negative correlation (Randazzo et al., 2021).

530 In the Massif Central, $\delta^{13}\text{C}$ varies between -6.3 ‰ and -3.6 ‰ while CO₂/³He
531 ranges from 3.9×10^9 to 1×10^{13} (Bräuer et al., 2017), which falls between upper mantle
532 and marine limestone in Fig. 6. This range can be interpreted as representing the
533 composition of the typical regional mantle end-member. In this case, we might observe,
534 for the SP4 group, the occurrence of a mantle-like component associated with CO₂
535 dissolution in water leading to more depleted $\delta^{13}\text{C}$ or a small contribution from organic
536 sediments, and at Fontaines Salées and Bourbon-Lancy, an organic source.
537 Alternatively, CO₂ in the SP4 group samples might originate from non-marine
538 carbonates or graphite which can range, respectively, from $\sim -18 \text{ ‰}$ to 0 ‰ and $\sim -40 \text{ ‰}$
539 to 0 ‰ (Favara et al., 2002) that can still be associated with mixing with a MORB
540 component. The history of CO₂ transport might be more complex for Fontaines Salées
541 and Bourbon-Lancy and could reflect secondary processes.

542 CO₂ reservoirs worldwide are generally interpreted as being magmatic in origin,
543 with $\delta^{13}\text{C}$ ranging between -5.7 to -2.7 ‰ (Gilfillan et al., 2009); compared to them
544 our dataset is atypical. However, in such reservoirs, water is the storage medium of CO₂
545 (Dubacq et al., 2012; Gilfillan et al., 2008) while in our study we did not identify water
546 bodies at depth and no water manifested during the well tests. Nonetheless, in areas
547 with no evidence of volcanism, CO₂ flow in the surface has already been reported for

548 seismically active areas (e.g. Li Vigni et al., 2022), indicating that fault systems can
549 lead to complex histories of CO₂ transport and entrapment.

550 For all samples, the $\delta^{15}\text{N}$ values show little variation, ranging from +6.1 to +8.4
551 ‰. This range is typical of continental or subduction-related gases and differs from that
552 of MORB (around -5 ‰ in the convecting mantle; Marty & Dauphas, 2003). Mantle
553 plumes also show lower $\delta^{15}\text{N}$ values, around 0 ± 3 ‰ (Labidi et al., 2020; Marty &
554 Dauphas, 2003). In the Massif Central, $\delta^{15}\text{N}$ values around +2 ‰ (Bräuer et al., 2017)
555 could potentially indicate the presence of a mantle plume or, alternatively, mixing
556 between MORB-like and continental-like components. The near-constant $\delta^{15}\text{N}$ values of
557 the present samples point to a common nitrogen source despite large variations in the N₂
558 contents of the surveyed springs. Such variable N₂ contents might be relative to
559 secondary physical/chemical processes in the crust since higher contents of N₂ and
560 helium are associated with more depleted $\delta^{13}\text{C}$ values (Fig S1).

561

562 **6 Conclusion**

563 The main source of helium in the study area is the continental crust. $^3\text{He}/^4\text{He}$
564 ratios are independent of gas composition, given that $^3\text{He}/^4\text{He}$ is roughly the same for
565 springs near SP4, including Fontaine des Vertus, while its helium and N₂ contents are
566 much higher than for the other springs of the group. Local helium production alone is
567 unlikely since the $^3\text{He}/^4\text{He}$ ratios (0.17-0.22 R_a) are in the same range across a 40 km
568 area (Pougues-les-Eaux and Couy). To verify if ^3He and ^4He are produced and stored in
569 the reservoir layers or if they are transported from a crustal flux, the U-Th-Li contents
570 of the host rocks of the sites investigated is required, and a better understanding of the
571 basin geometry and sources is necessary.

572 The decreasing $^3\text{He}/^4\text{He}$ gradient starting at the Massif Central towards the Paris
573 Basin and the vertical $^3\text{He}/^4\text{He}$ gradient previously described in the literature at the Paris
574 Basin aquifers (Marty et al., 1993; Pinti & Marty, 1995, 1998) point to a small mantle
575 contribution limited to the N-S fault system, since laterally, few kilometres (50-80) east
576 at Bourbon-Lancy and Fontaines Salées, $^3\text{He}/^4\text{He}$ ratios show typical crustal values (\leq
577 0.04 R_a). The ^{129}Xe excess observed at SP7 confirms that excess ^3He relative to the
578 crustal end-member composition is derived from the mantle. This indicates that either a

579 deep-seated magmatic source is present, or deep fault systems reach the mantle and
580 input ^3He via an improved permeability system (Kennedy & van Soest, 2007).

581 Considering that the $\delta^{13}\text{C}$ of CO_2 of our surveyed springs are distinct from the
582 regional mantle value (Massif Central), CO_2 is probably mainly of crustal origin.
583 However, mixing between crustal or biogenic components and mantle is expected,
584 especially since CO_2 would most likely be the carrier phase of the mantle helium
585 component and $\text{CO}_2/{}^3\text{He}$ ratios (between 8×10^9 and 1×10^{10}) are close to the MORB
586 value and indicate a $\sim 15\%$ mantle contribution, in agreement with what is observed in
587 He and Xe isotopes. Secondary processes in the crust are also possible given the low
588 solubility of CO_2 in water, and potential carbonate precipitation can also explain the
589 decrease of CO_2 contents associated with elevated concentrations of N_2 and helium at
590 Fontaine des Vertus, Bourbon-Lancy, and Fontaines Salées. $\delta^{15}\text{N}$ results, for both
591 sample groups, are comparable and point to a common crustal source.

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Fig 1

Geological map of the study area with helium isotope results (in R_a) and literature data. a) Structural map with major structures inherited from the Variscan orogeny (Baptiste, 2016). b) Simplified geological map (1/1,000,000 BRGM French geological map) with main geological structures of the area (Au: Aumance basin, De: Decize horst and Ma: Machine graben), location of the well SP4, SP7, springs and the lithium deposits of the area (Ch: Chanvence and Bv: Beauvoir; Gloaguen et al., 2018), and distribution of literature helium isotope data (Bräuer et al., 2017; Marty et al., 1993; Pinti & Marty, 1998). c) Focus on the Saint-Parize-le-Châtel area with location of the springs and wells. d) Simplified cross section line of the study area, with location on the map “b)” marked by the orange line.

Fig 2

$^4\text{He}/^{20}\text{Ne}$ versus $^3\text{He}/^4\text{He}$ plot. Samples plot in two groups, one with signatures close to a crustal source and the other with either a 2.5% input of mantle or with nucleogenic ^3He production consistent with the presence of 600 ppm Li. Lithium contents for the lower and upper limits of the European continental crust range are based on average values of U-Th for the continental crust (1.8 and 7.2 ppm, respectively; Krauskopf and Bird, 1967). $^4\text{He}/^{20}\text{Ne}$ values for SP4 are not precise due to ^{20}Ne abundances indistinguishable from baseline measurement. Massif Central and Paris Basin data are from the literature (Bräuer et al., 2017; Marty et al., 1993; Pinti & Marty, 1998).

Fig 3

Variation of helium isotope ratios with latitude. There is an apparent trend with decreasing $^3\text{He}/^4\text{He}$ ratios from south, starting at the Massif Central, to north, ending in the Paris Basin. Helium contents are higher for our dataset. Massif Central and Paris Basin data are from the literature (Bräuer et al., 2017; Marty et al., 1993; Pinti & Marty, 1998).

Fig 4

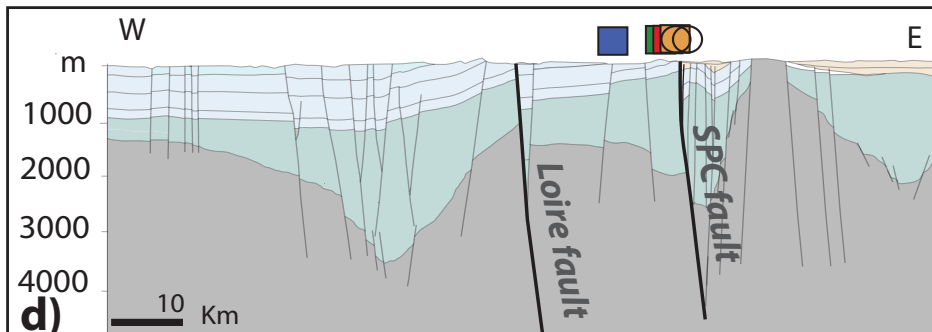
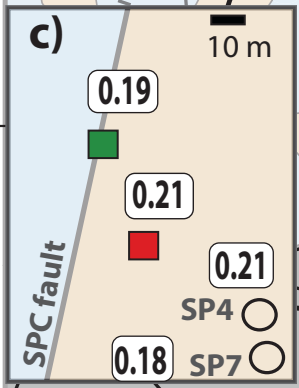
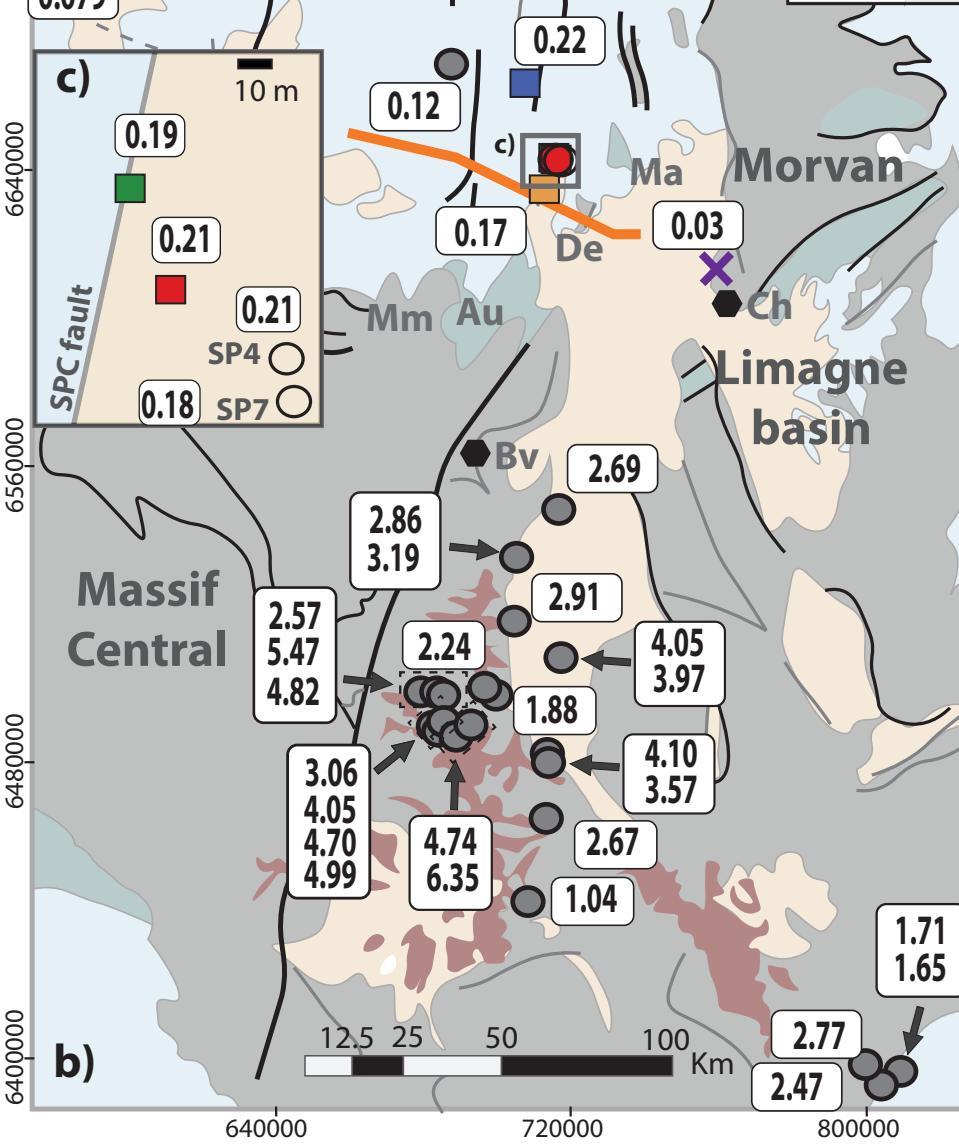
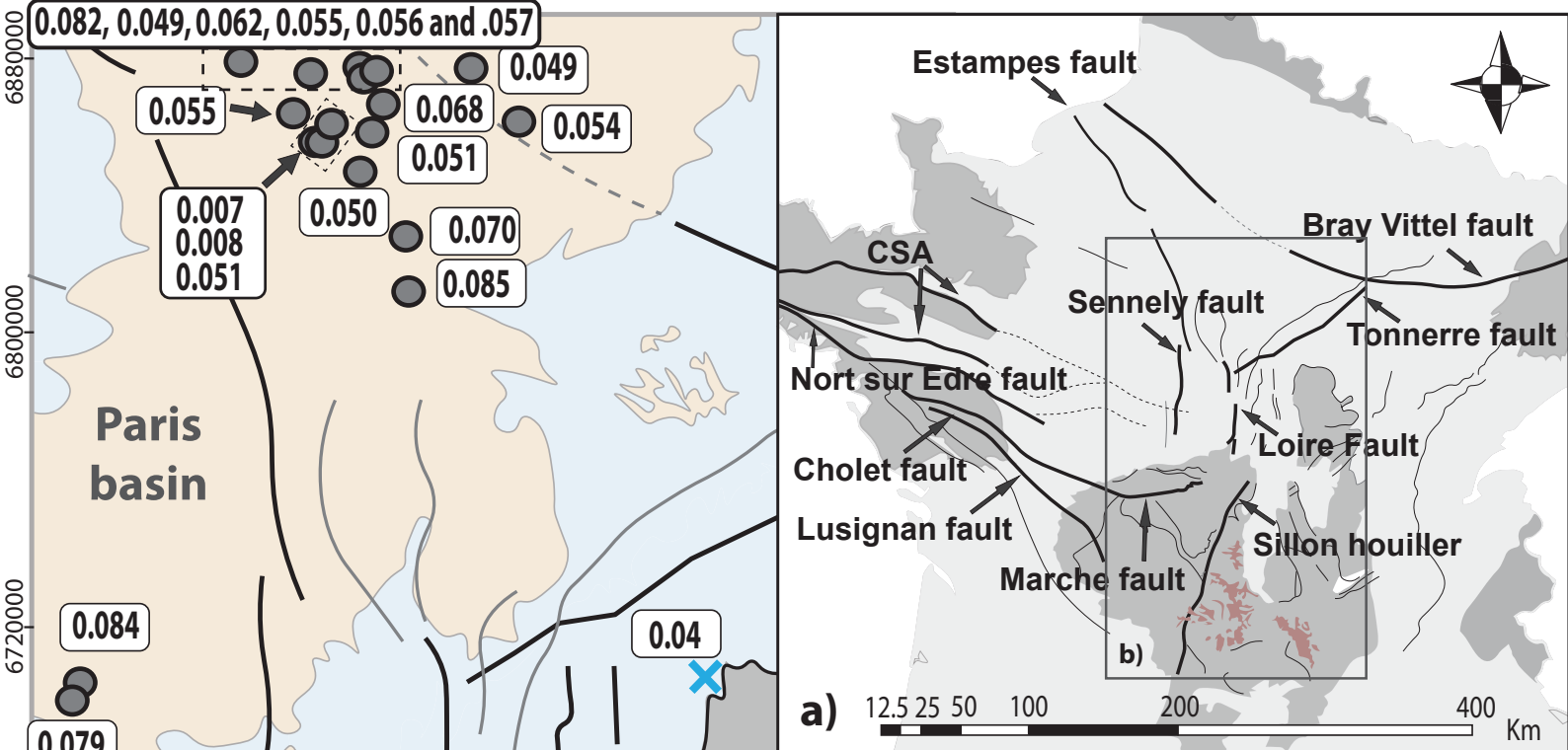
Three isotope diagram of neon. Mixing between air and solar, MORB (Sarda et al., 1988), Archean terrains (Lippmann-Pipke et al., 2011), crustal (Kennedy et al., 1990) end-members are displayed. SP4 plots between the crustal and Archean mixing lines with air and errors shown in 1 sigma level.

Fig 5

Xenon isotopic composition of SP7 gas corrected for physical fractionation. Data is expressed in δ (‰) of Xe ratios relative to atmospheric composition and normalised to ^{130}Xe . Error bars are 2 sigma. The SP7 gas has a clear ^{129}Xe excess that indicates mantle input.

Fig 6

$\text{CO}_2/^3\text{He}$ versus $\delta^{13}\text{C}$ plot (Sano and Marty, 1995). The dataset shows mixing between the three components and plots close to the MORB value while Massif Central data (Bräuer et al., 2017) only shows mixing between MORB and marine limestone. The grey field represents the MORB range from Marty & Zimmermann (1999).



Geological units

Sedimentary cover

- Cenozoic
- Mezosoic
- Stephanian - Permian

Basin

- Paris basin
- Limagne basin

Basement

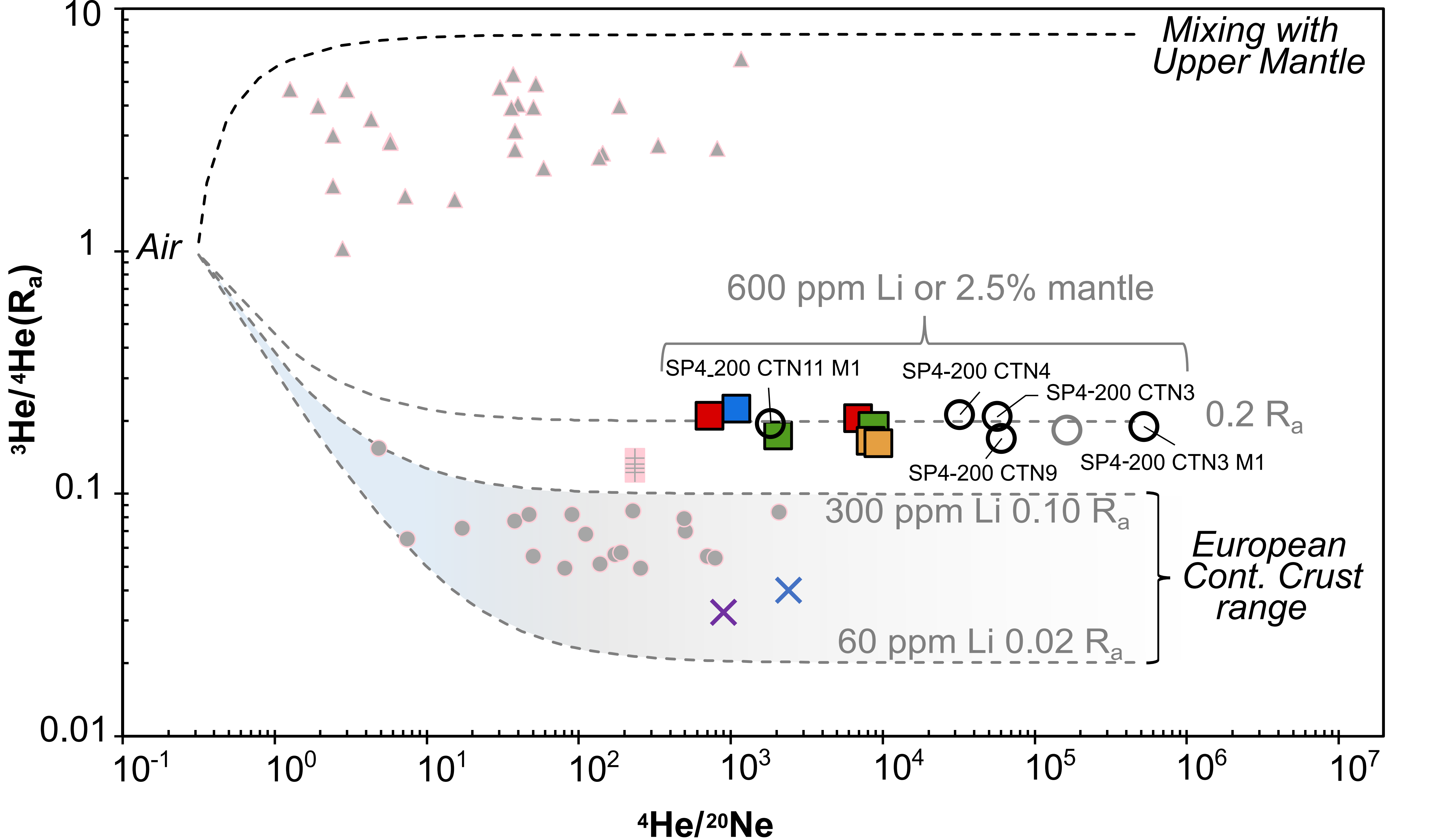
- Magmatic, metamorphic and volcanic units
- Tertiary volcanism
- Lithium deposit

Structural elements

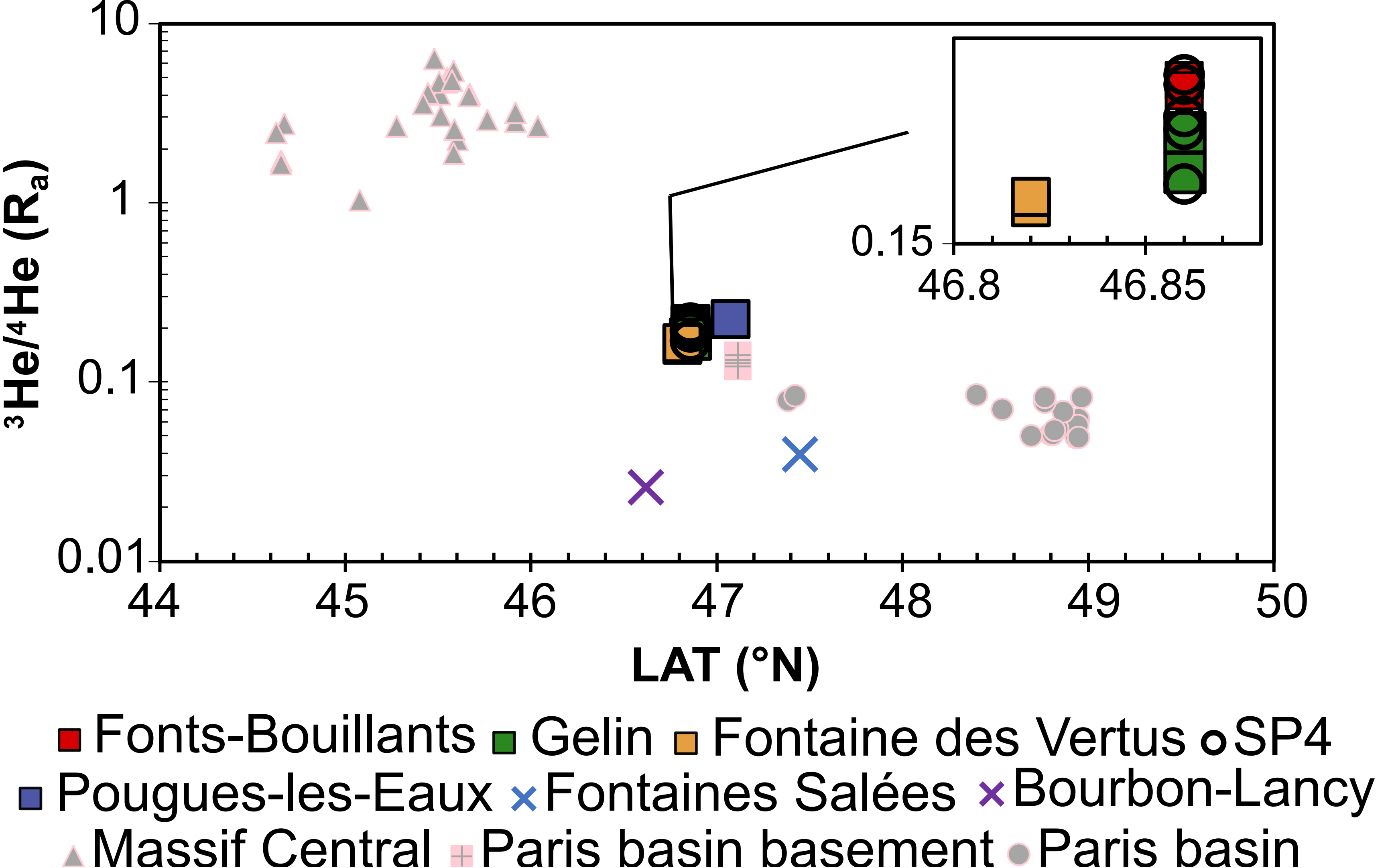
- SPC Saint-Parize-le-Châtel
- Faults
- - Expected faults
- Main faults
- Cross section line

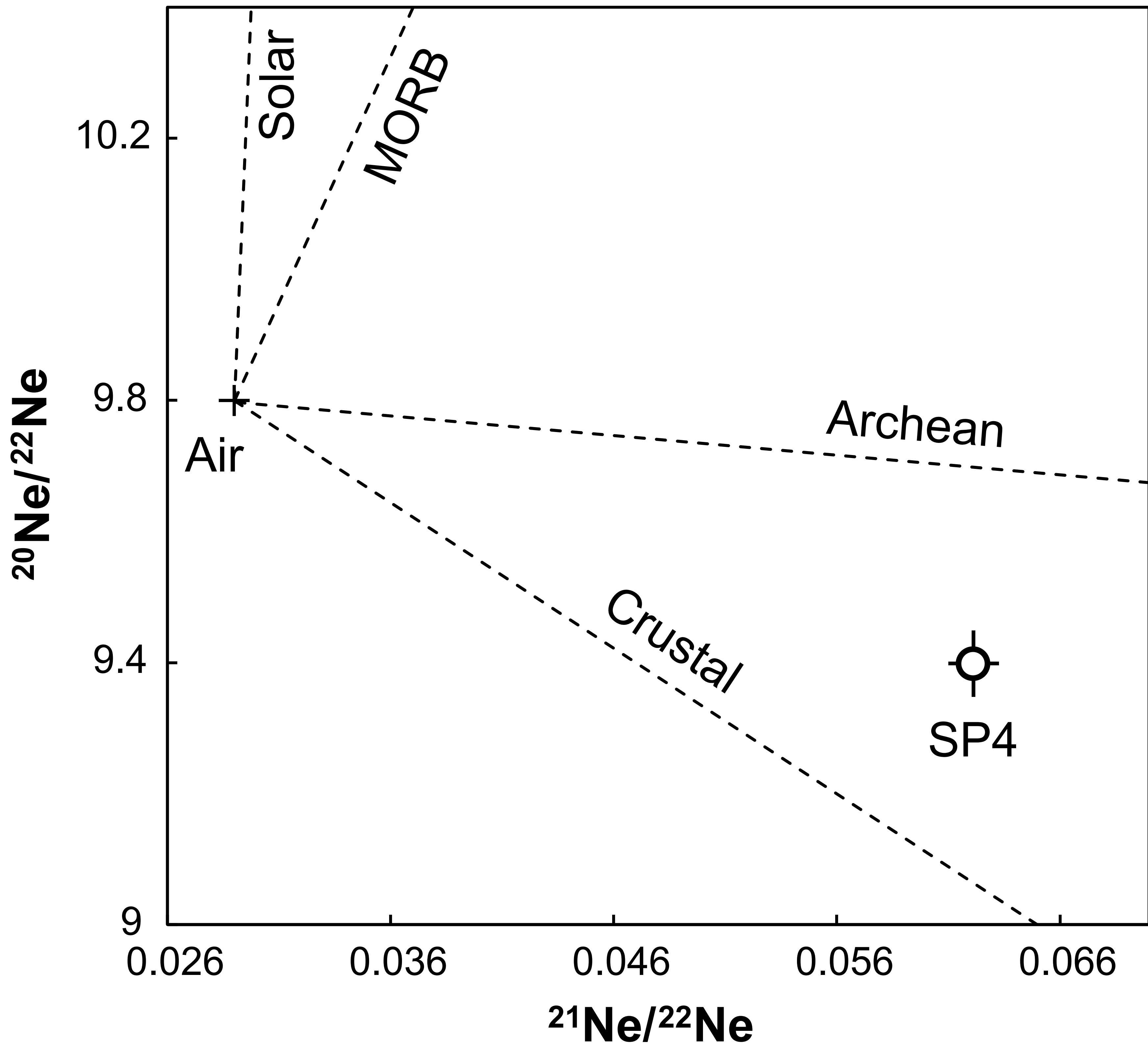
Springs and wells

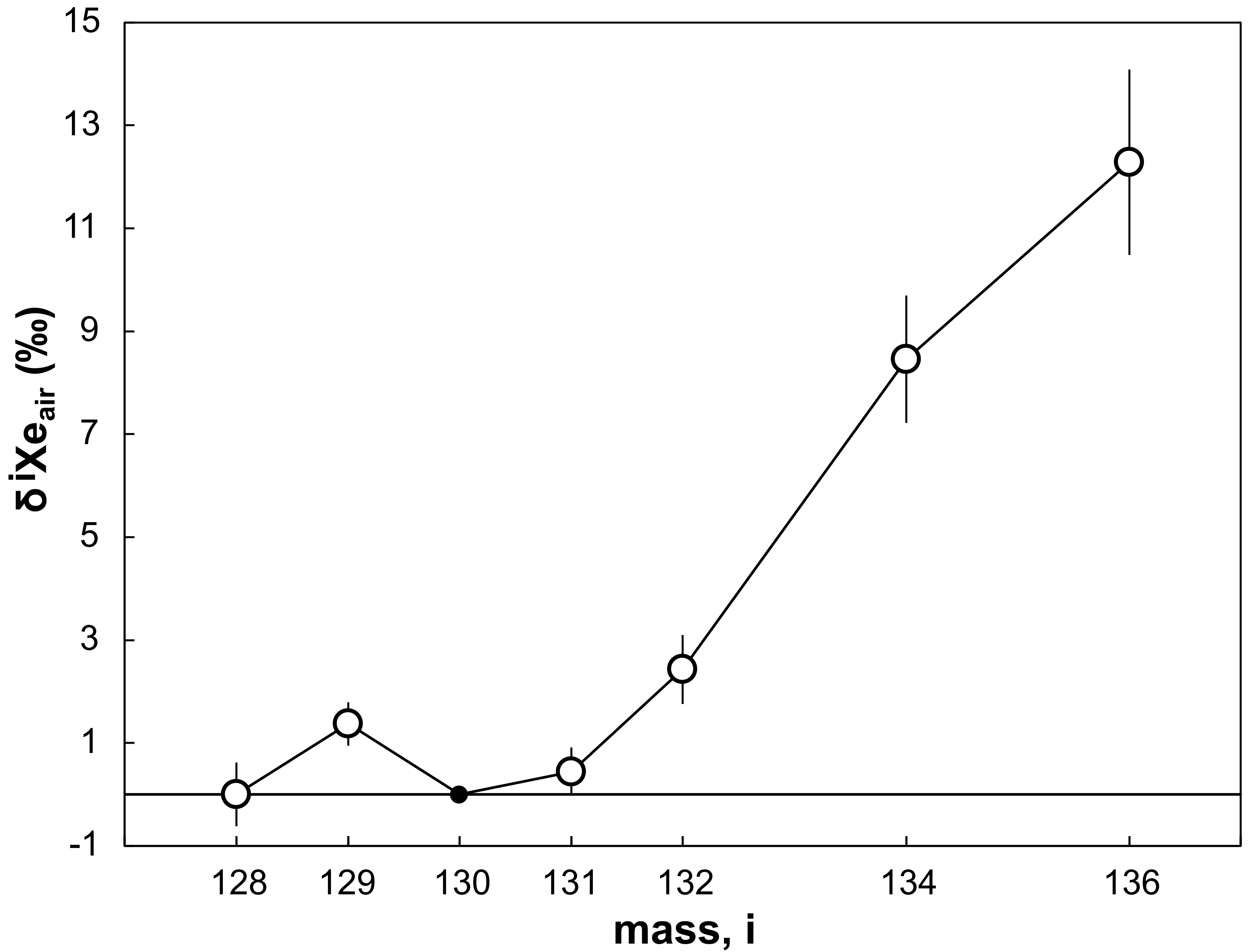
- Wells and springs from litterature
- Gelin
- Pougues-les-Eaux
- Fontaines-Bouillants
- Fontaines des Vertus
- SP4 and SP7 wells
- Fontaines Salées
- Bourbon-Lancy
- R/Ra ratio



■ Fonts-Bouillants ■ Gelin ■ Fontaine des Vertus ■ Pougues-les-Eaux × Fontaines Salées
 × Bourbon-Lancy ○ SP4 ○ SP7 ▲ Massif Central ■ Paris basin basement ● Paris basin







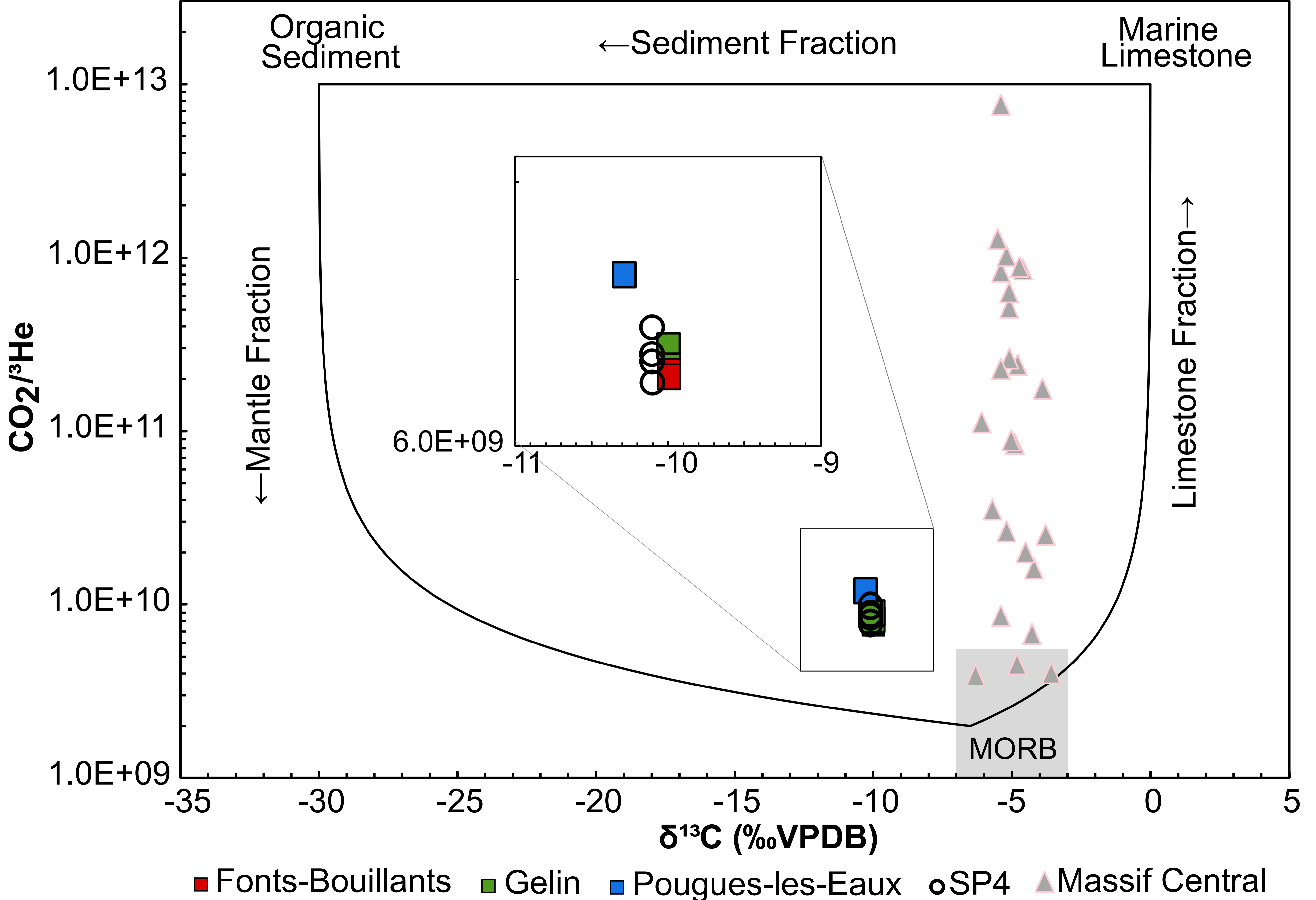


Table 1 List of **sampled sites** with coordinates, **sampling dates**, and **average of main** gas phase results from GC measurements. ***A second sampling campaign occurred in May-June/2022 for $\delta^{15}\text{N}$ analyses.**

Site	type	Date	map ref	LAT (°N)	LONG (°E)	CO ₂ (vol.%)	N ₂ (vol.%)	He (ppmV)
SP4	well	Aug-Sept/2022	1	46.85979	3.19336	93	5.6	394
SP7	well	Feb/2022	2	47.82861	4.18083	94	5.6	426
Gelin	spring	Aug-Sept/2022	3	46.86035	3.19268	91	6.1	416
Fonts-Bouillants	spring	Aug-Sept/2022	4	46.86008	3.19284	92	5.7	393
Fontaine des Vertus	spring	Aug-Sept/2022*	5	46.82025	3.17123	60	36.3	2874
Pougues-les-Eaux	spring	Aug-Sept/2022*	6	47.07635	3.09240	94	4	245
Fontaines Salées	spring	Aug-Sept/2022	7	47.44900	3.77680	0.9	92.5	45794
Bourbon-Lancy	spring	May-June/2022	8	46.61786	3.76974	2.6	91.3	20644

Table 2 $^3\text{He}/^4\text{He}$, $^4\text{He}/^{20}\text{Ne}$, $\text{CO}_2/{}^3\text{He}$, and stable isotopes results. $^3\text{He}/^4\text{He}$ are normalized by the atmospheric $^3\text{He}/^4\text{He}$ ratio ($R_a = 1.39 \times 10^{-6}$; Graham, 2002). *Average $\delta^{15}\text{N}$ values from multiple measurements; SP4 value from two measurements performed on the same day (6.3 and 8.3 ‰) and Gelin from three measurements on different days (7.9, 8.5, and 8.9 ‰).

Site	Sample	$^4\text{He}/^{20}\text{Ne}$	$^3\text{He}/^4\text{He}$ (R_a)	$\text{CO}_2/{}^3\text{He}$	$\delta^{13}\text{C}$ (‰ vs PDB)	$\delta^{15}\text{N}$ (‰ vs Air)
SP4	SP4-100 CTN6	-	-	-	-10.1	7.3*
	SP4-200 CTN3	62080 ± 4724	0.21 ± 0.01	8.16E+09	-	-
	SP4-200 CTN3 M1	≤ 1656612	0.19 ± 0.01	8.96E+09	-	-
	SP4-200 CTN4	35291 ± 3003	0.21 ± 0.00	7.99E+09	-	-
	SP4-200 CTN9	66651 ± 5465	0.17 ± 0.01	1.00E+10	-	-
	SP4-200 CTN11 M1	1975 ± 55	0.20 ± 0.01	8.71E+09	-	-
SP7	SP7	180114 ± 34276	0.18 ± 0.01	8.66E+09	-	-
Gelin	GEL-100 CTN1	9509 ± 354	0.19 ± 0.00	8.37E+09	-10.0	8.4*
	GEL-300 CTN1 M1	2222 ± 72	0.17 ± 0.01	9.09E+09	-	-
Fonts-Bouillants	FBO-100 CTN1 M1	776 ± 22	0.21 ± 0.00	8.00E+09	-10.0	-
	FBO-300 CTN1 M1	7525 ± 281	0.21 ± 0.01	8.18E+09	-	-
Fontaine des Vertus	VER-100 CTN1	10240 ± 336	0.16 ± 0.01	9.26E+08	-	-
	VER-200 CTN1 M1	9042 ± 476	0.17 ± 0.00	9.06E+08	-11.8	7.5
Pougues-les-Eaux	POU-100 CTN1 M1	1170 ± 35	0.22 ± 0.01	1.22E+10	-10.3	6.1
Fontaines Salées	FON-100 CTN1 M1	2587 ± 66	0.04 ± 0.00	-	-20.0	7.1
Bourbon-Lancy	B-LANCY A	960 ± 20	0.03 ± 0.01	-	-	-

Table 3 Ne, Ar, and Xe isotope results. Xe ratios are reported in the delta notation relative to the atmospheric isotope composition. “corr” refers to data corrected for physical fractionation. Errors shown at 1 sigma.

Sample	$^{20}\text{Ne}/^{22}\text{Ne}$	$^{21}\text{Ne}/^{22}\text{Ne}$	$^{40}\text{Ar}/^{36}\text{Ar}$	$^{38}\text{Ar}/^{36}\text{Ar}$	δ vs Air, ‰					
					$^{128}\text{Xe}/^{130}\text{Xe}$	$^{129}\text{Xe}/^{130}\text{Xe}$	$^{131}\text{Xe}/^{130}\text{Xe}$	$^{132}\text{Xe}/^{130}\text{Xe}$	$^{134}\text{Xe}/^{130}\text{Xe}$	$^{136}\text{Xe}/^{130}\text{Xe}$
SP4-200-GB1	9.40 ± 0.05	0.0621 ± 0.0011	2866 ± 5	0.189 ± 0.001	-	-	-	-	-	-
SP4-100-GB1	-	-	2574 ± 11	0.179 ± 0.001	-	-	-	-	-	-
SP7 (raw)	-	-	696.6 ± 0.2	0.18840 ± 0.00002	0.996 ± 0.311	1.865 ± 0.145	-0.042 ± 0.179	1.463 ± 0.145	6.554 ± 0.167	9.464 ± 0.195
SP7 (corr)	-	-	-	-	0 ± 0.311	1.371 ± 0.212	0.445 ± 0.235	2.430 ± 0.335	8.462 ± 0.619	12.286 ± 0.902